

# Stochastic field lines and braids

Jean-Luc Thiffeault

Department of Mathematics  
University of Wisconsin – Madison

Plasma Theory Seminar, University of Wisconsin, Madison  
10 December 2012

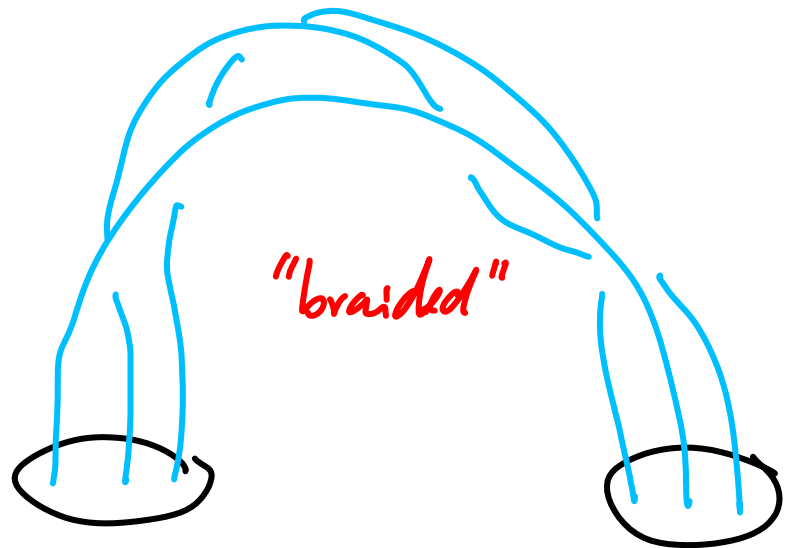
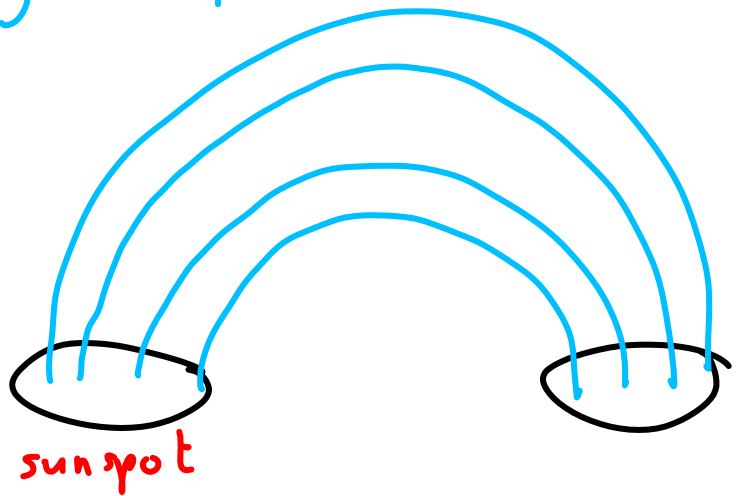
Supported by NSF grants DMS-0806821 and CMMI-1233935



# Tangled magnetic fields

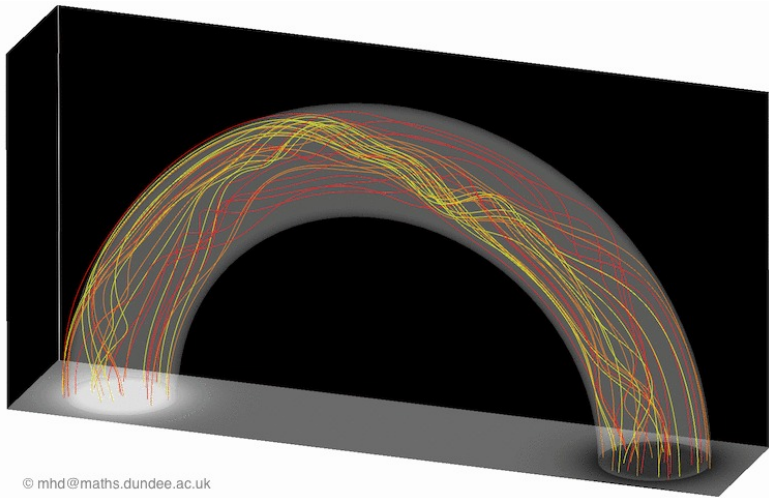
The "vision" for solar flux tubes

magnetic field



The magnetic field lines become "braided" due to MHD frozen-in condition + turbulence

## Tangled magnetic fields (cont'd)



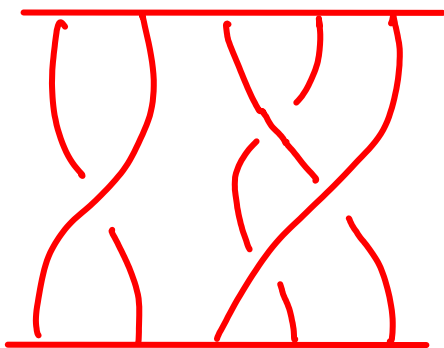
© mhd@maths.dundee.ac.uk

Source: <http://www.maths.dundee.ac.uk/mhd/>



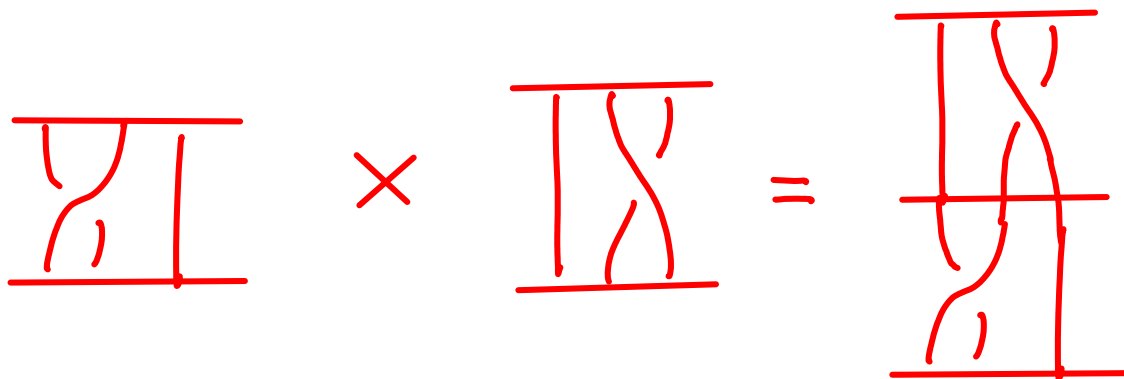
# The braid group $B_n$

A braid is a set of  $n$  strands with fixed endpoints

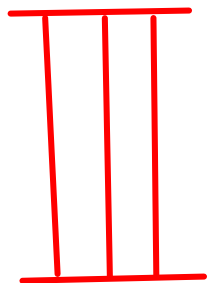


Two braids are equal if they can be "deformed" into each other, whilst holding ends fixed. "ambient isotopy"

Braids form a group:  
(for fixed  $n$ )

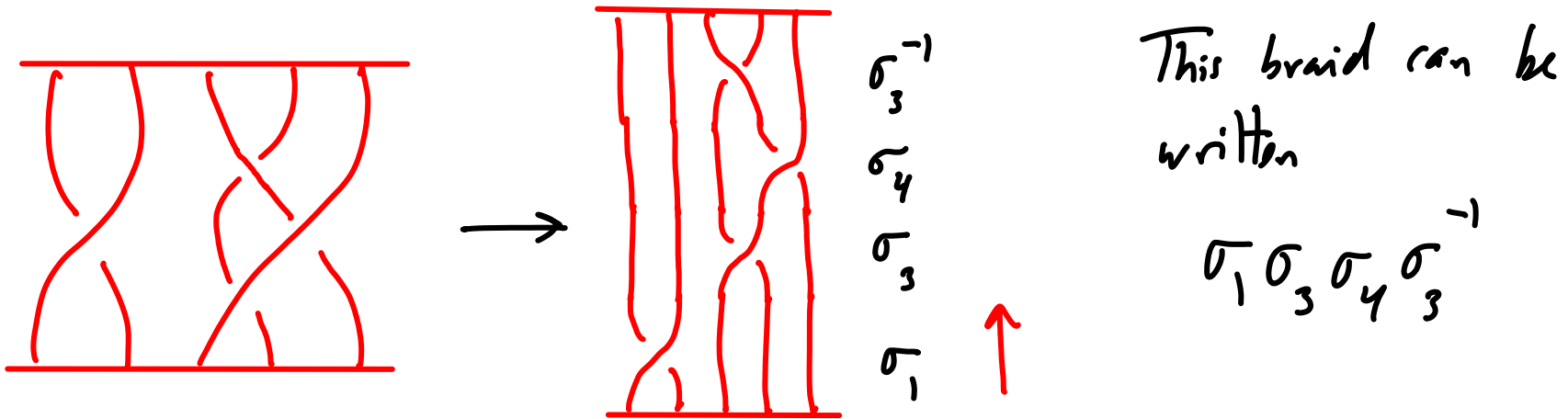


Identity:



This is associative, and for every braid there is an inverse that "disentangles" the braid.

# Braid generators



$\{\sigma_1, \sigma_2, \dots, \sigma_{n-1}\}$  are generators of  $B_n$

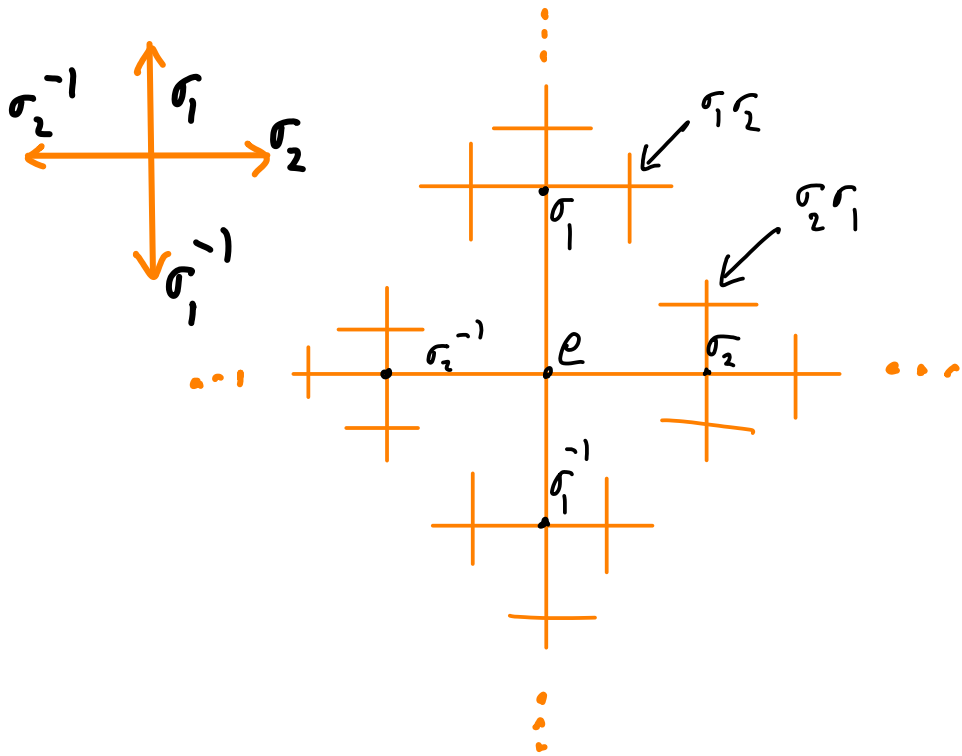
They satisfy relations:  $\sigma_i \sigma_j = \sigma_j \sigma_i$ ,  $|i-j| > 1$

$$\sigma_i \sigma_j \sigma_i = \sigma_j \sigma_i \sigma_j, \quad |i-j| = 1$$

**Artin** proved that these are the only relations that arise from physical braids

# Cayley graph

A convenient graphical representation of groups is as a graph:



The Cayley graph for  $B_3$  might start out like this, but there are "shortcuts" (loops) due to braid relations.

Now we can define a random walk on this graph by choosing a random direction to move to from each vertex.

$$\gamma = \text{random word} = \sigma_1 \sigma_2 \sigma_1^{-1} \sigma_1 \sigma_2^{-1} \sigma_1 \sigma_1 \dots \quad (\text{say})$$

Lots of interesting questions! (recurrence, distance, etc...)

The simplest Cayley graph

For  $B_2$ , only one generator



The exponent gives the winding number of one strand around the other.

If our random walk moves to the left/right with probability  $p/(1-p)$  expect net winding number  $m$  after  $N$  steps to have probability

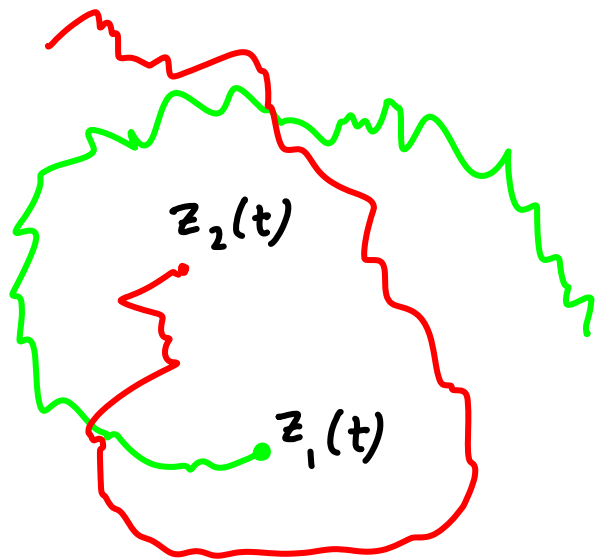
$$P(m) = \binom{N}{k_m} p^{k_m} (1-p)^{N-k_m} \quad k_m = \frac{1}{2}(m+N)$$

( $m+N$  even!)

Mean  $N(p-1/2)$ , variance  $Np(1-p) \Rightarrow$  Gaussian for large  $N$



# Winding number of two Brownian processes



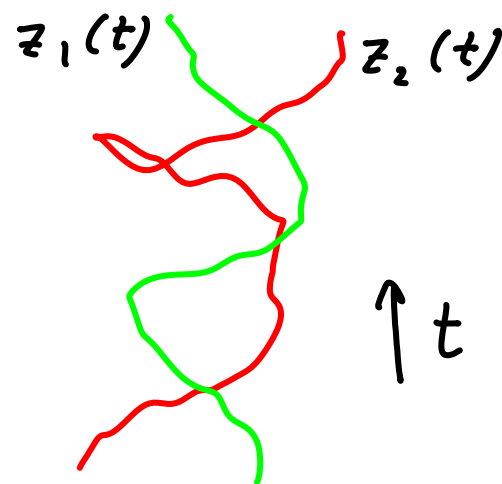
What is the distribution of the winding angle,  $\theta$ , after a large time  $t$ ?

Same as Cayley graph random walk?

Consider now two Brownian processes on the plane, diffusivity  $D$ .  $\langle z_i^2 \rangle \sim 2Dt$

(Think of two stochastic field lines)

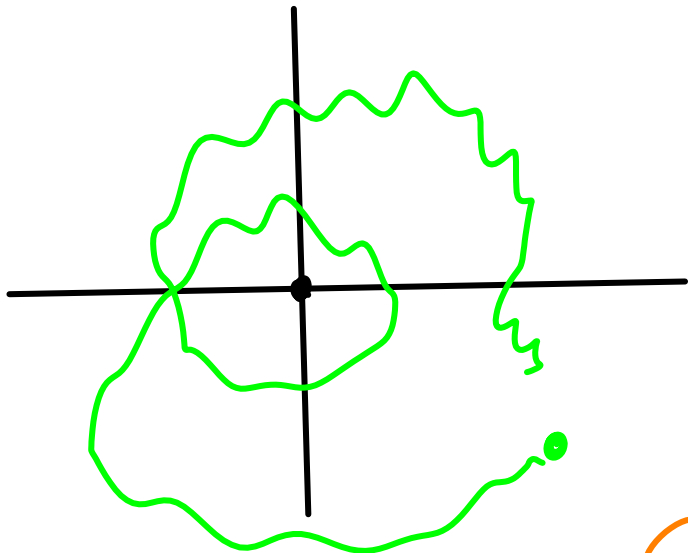
Picture as a braid:



## Winding around the origin

Let  $l(t) = z_1(t) - z_2(t) \rightarrow$  Brownian with diffusivity  $2D$   
 $\langle l^2 \rangle = \langle z_1^2 \rangle + \langle z_2^2 \rangle = 2(2Dt)$

So now the question is: how many times does  $l$  wind around the origin?



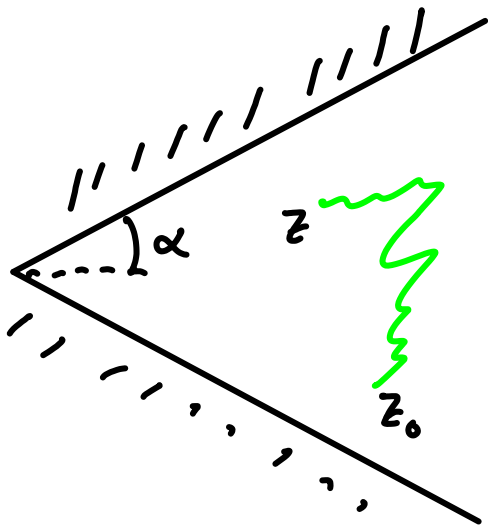
Classic result of Spitzer (1958):

$$p(x) \sim \frac{1}{\pi} \frac{1}{1+x^2}$$

$$x = 2\theta / \log(4Dt/r_0^2) \quad \frac{Dt}{r_0^2} \gg 1$$

Cauchy-Lorentz  $\rightarrow$  not Gaussian!

How do we show this?



Brownian process  $\Rightarrow$  heat equation

Solve in wedge of angle  $2\alpha$ :

$$\frac{\partial P}{\partial t} = D \nabla^2 P \quad P_0 = \delta(z - z_0)$$

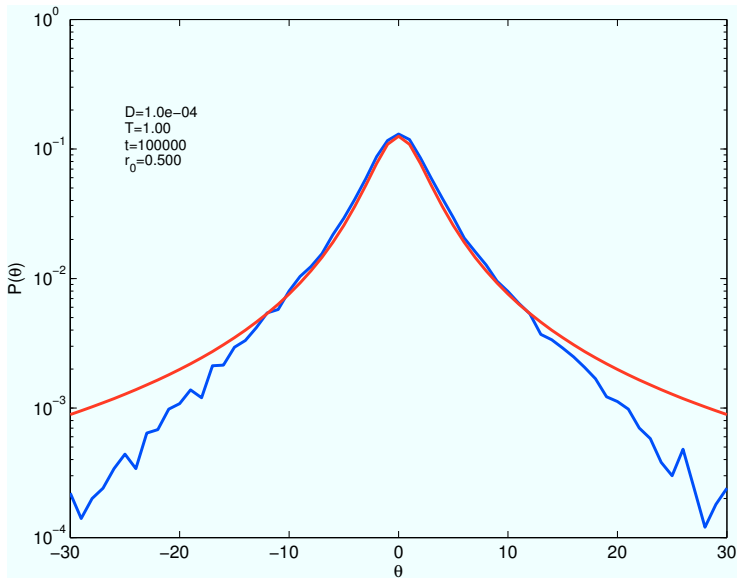
with Neumann conditions (conservation of probability)

Compute Green's function (e.g. Carslaw & Jaeger)

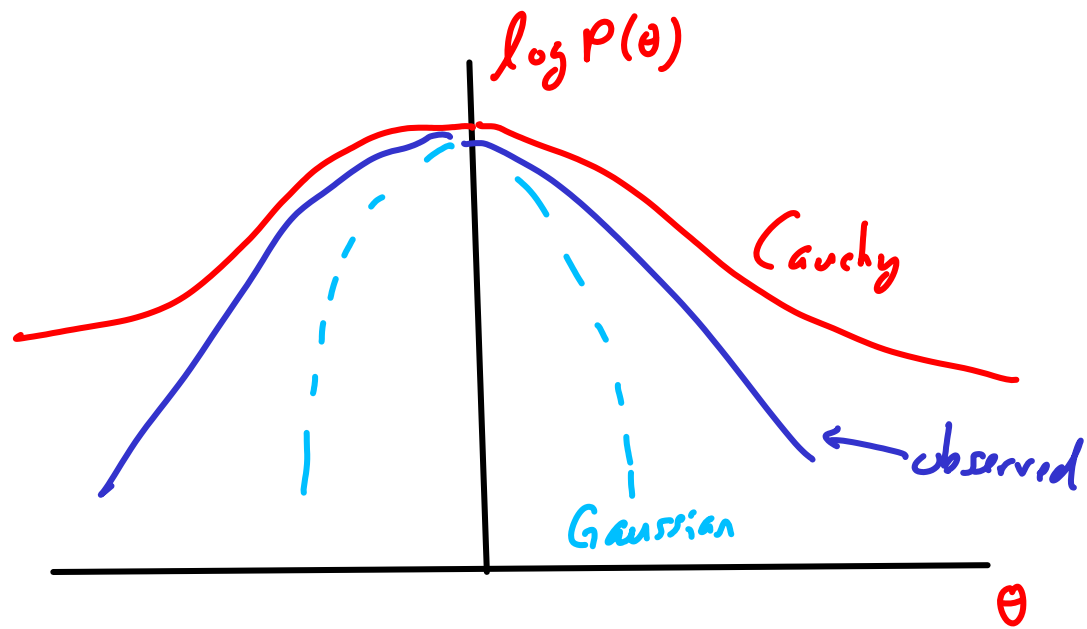
then take  $\alpha \rightarrow \infty!$   $\rightarrow$  multiple Riemann sheets

large  $-t$  asymptotics then give Cauchy distribution.

# Numerical simulations



# The problem with the tails



numerical sims.  
In practice, we never see this. The large winding angles predicted by Spitzer are a symptom of the scale-free Brownian process.  
→ can wind very fast around origin.

Any "regularization" (random walk, length scale, curvature-limited...) gives

$$p(x) \sim \frac{1}{2} \operatorname{sech}\left(\frac{\pi x}{2}\right)$$

← exponential tails  
(still not Gaussian)

[One way to get this: take out disk around origin]

[Pitman & Yor '86, Berger '87, Drosuel & Kardar '96, Grosberg & Frisch '03]

## Scaling of $x$ with $\log t$

The scaling  $x \sim 2\theta / \log(4Dt/r_0^2)$  arises because the Brownian process wanders away from the origin  $\rightarrow$  PDF stops changing

Scaling argument:

$$\frac{dr}{d\theta} = r f(\theta) \quad \text{since } r \text{ is the only length scale for } Dt \gg r_0^2$$

$\hookrightarrow$  indep. of  $\theta$  by isotropy

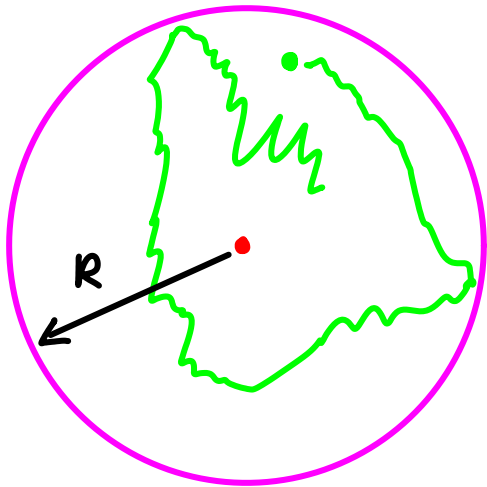
$$\text{So } d\theta \sim \frac{dr}{r} \sim \frac{dt}{t} \sim d \log t \quad \text{since } r \sim t^{1/2}$$

[See Fisher et al. 1984, Drossel & Kardar 1996]

## Closed domain

Now all this was for a Brownian motion on the plane.

In a confined geometry, the process "starts over" when it reflects. [Markovian]



So pieces of random walk of duration  $\frac{R^2}{D}$  diffusion time across disk

are uncorrelated, and each piece has a Spitzer distribution.

Convolving the  $t/(R^2/D)$  distributions then gives

$$x \sim \frac{\theta}{\sqrt{t}}$$

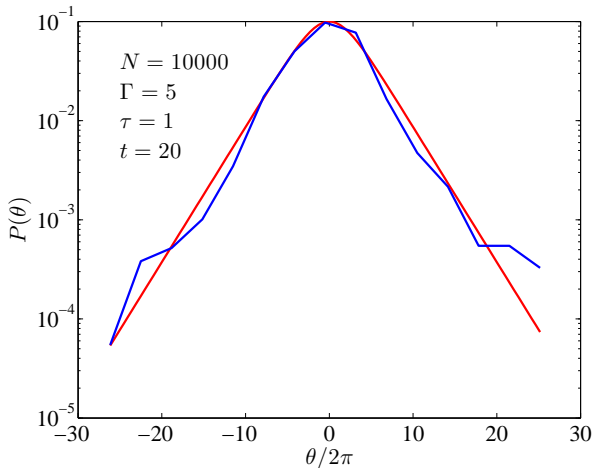
rather than  $\theta/\log t$

[Drossel & Kardar 2003]

Confinement leads to many more turns

## Blinking vortex simulations

Winding number for blinking vortex pair in a disk (Aref, 1984):



The red curve is the sech distribution, with fitted diffusivity.



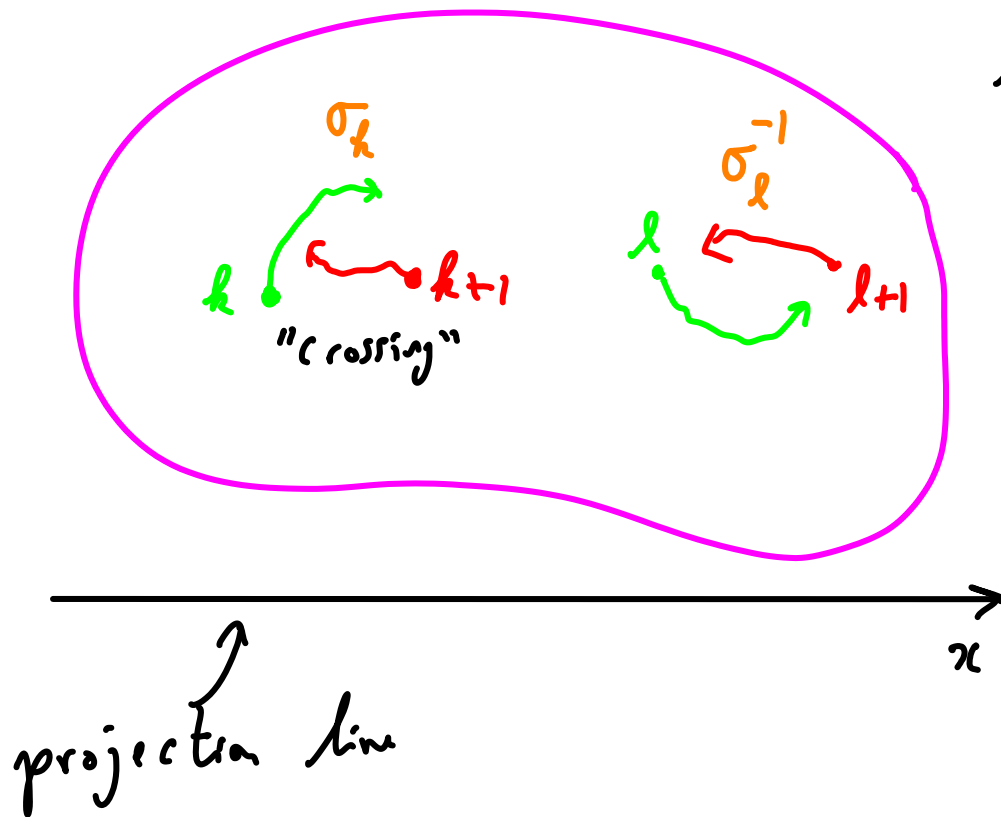
## The moral so far

Maybe I've convinced you that the random walk on a Cayley graph (which suggest Gaussian winding angle) is hard to reconcile with the twisting of two random walkers. (Not sure how to do it...)

At least in our two-walker example the probability of winding in one direction was the same as the other.  $P(\sigma_i) = P(\sigma_i^{-1})$

But more generally, for  $n$  random walks, are all braid generators  $\{\sigma_1, \dots, \sigma_{n-1}\}$  equally likely?

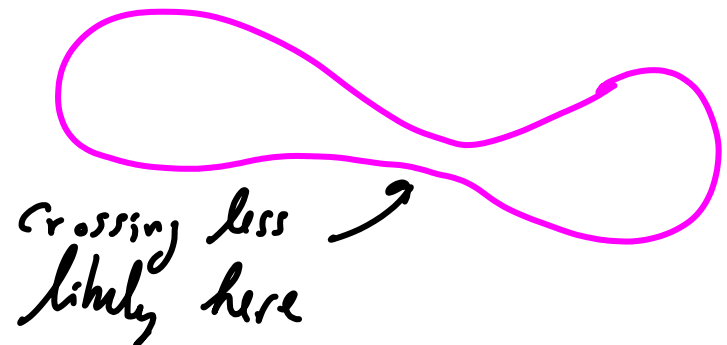
# Random walkers in an arbitrary domain



When projecting along a fixed line, each "crossing" of two walkers corresponds to a generator  $\sigma_k^{\pm 1}$ .

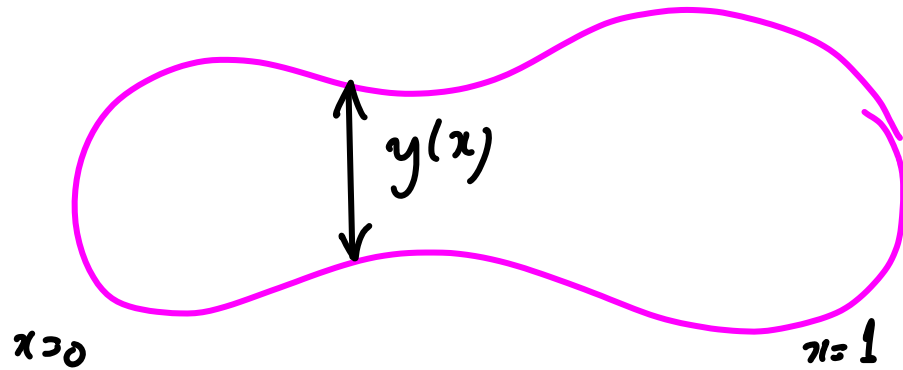
Since each walker is independent and uniformly distributed, the probability of observing a crossing depends on width:

[See JLT 2005, 2010]



Crossing less likely here

Where do crossings occur?



The probability of a walker being in  $[x, x+dx)$  is

$$\rho(x) dx = \frac{1}{A} y(x) dx$$

(So  $\int_0^1 \rho(x) dx = 1$ )  $\leftarrow$  area

The probability of a crossing occurring at  $x$  is

$$P(\text{crossing at } x) = \frac{\rho^2(x)}{\int_0^1 \rho^2(x) dx} \leftarrow \text{two particles at } x \text{ is a crossing}$$

[This is in the small step-size limit]

Which generator?

Once we have a crossing, we can ask which generator it corresponds to.

This depends on the ordering of the particles.

Find: [See Sarah Tumas's thesis]

$$P(\sigma_h^{\pm 1}) = \binom{n-2}{h-1} \int_0^1 p^2(x) P^{h-1}(x' < x) P^{n-h-1}(x' > x) dx$$

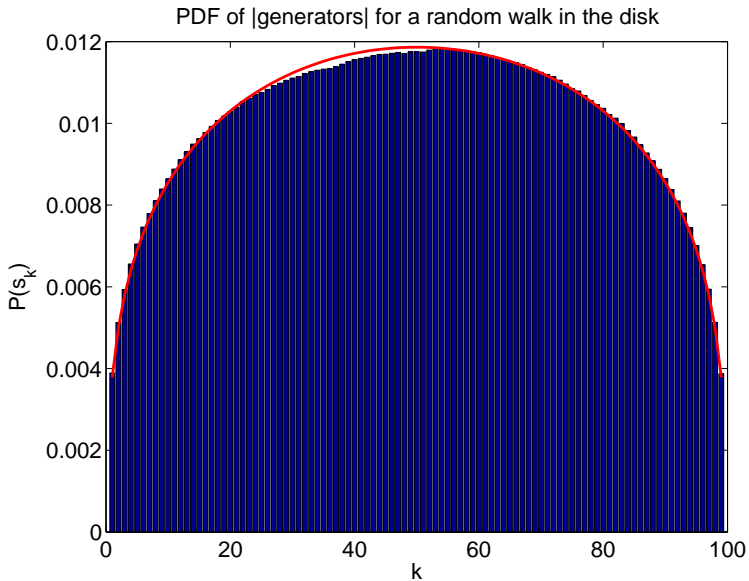
$\uparrow$  crossing at  $x$        $\uparrow$  particles to the left of  $h$        $\uparrow$  particles to the right of  $h$

For a square domain,  $p(x) = 1$  and

$$P(\sigma_h^{\pm 1}) = \frac{1}{n-1}$$

All generators projected along the  $x$ -axis occur with equal probability.

# Distribution of generators in a disk



# Conclusions

- Can generate braids by 'randomly picking generators' (**random walk on Cayley graph**), but not clear what physical process that corresponds to;
- Brownian motions have **Cauchy-distributed** winding angles;
- Random walks have **sech-distributed** winding angles, as does a simple chaotic flow;
- The braids created by random walks **depend on the shape of the domain!**
- Topological entropy of random braids?

# References

- Aref, H. 1984 Stirring by Chaotic Advection. *J. Fluid Mech.* **143**, 1–21.
- Bélisle, C. 1989 Windings of random walks. *Ann. Prob.* **17**, 1377–1402.
- Bélisle, C. & Faraway, J. 1991 Winding angle and maximum winding angle of the two-dimensional random walk. *J. Appl. Prob.* **28**, 717–726.
- Berger, M. A. 1987 The random walk winding number problem: convergence to a diffusion process with excluded area. *J. Phys. A* **20**, 5949–5960.
- Berger, M. A. & Roberts, P. H. 1988 On the winding number problem with finite steps. *Adv. Appl. Prob.* **20**, 261–274.
- Drossel, B. & Kardar, M. 1996 Winding angle distributions for random walks and flux lines. *Phys. Rev. E* **53**, 5861–5871.
- Fisher, M. E., Privman, V. & Redner, S. 1984 The winding angle of planar self-avoiding walks. *J. Phys. A* **17**, L569.
- Grosberg, A. & Frisch, H. 2003 Winding angle distribution for planar random walk, polymer ring entangled with an obstacle, and all that: Spitzer-Edwards-Prager-Frisch model revisited. *J. Phys. A* **36**, 8955–8981.
- Nechaev, S. K. 1996 *Statistics of Knots and Entangled Random Walks*. Singapore; London: World Scientific.
- Pitman, J. & Yor, M. 1986 Asymptotic laws of planar Brownian motion. *Ann. Prob.* **14**, 733–779.
- Pitman, J. & Yor, M. 1989 Further asymptotic laws of planar Brownian motion. *Ann. Prob.* **17**, 965–1011.
- Rudnick, J. & Hu, Y. 1987 The winding angle distribution for an ordinary random walk. *J. Phys. A* **20**, 4421–4438.
- Spitzer, F. 1958 Some theorems concerning 2-dimensional Brownian motion. *Trans. Amer. Math. Soc.* **87**, 187–197.
- Summers, D. W. 2009 Random knotting: Theorems, simulations and applications. In R. L. Ricca, editor, *Lectures on Topological Fluid Mechanics*, pp. 187–217. Berlin: Springer.
- Thiffeault, J.-L. 2005 Measuring Topological Chaos. *Phys. Rev. Lett.* **94**, 084502.
- Thiffeault, J.-L. 2010 Braids of entangled particle trajectories. *Chaos* **20**, 017516. [arXiv:0906.3647](https://arxiv.org/abs/0906.3647).
- Tumasz, S. E. 2012 *Topological Stirring*. Ph.D. thesis, University of Wisconsin – Madison, Madison, WI.