Nonlinear MHD Stability and Dynamical Accessibility

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Overview

- We discuss a unified description of variational methods for establishing stability of plasma equilibria.
- The first method is based upon a Lagrangian approach (in the sense of fluid elements). A Lagrangian equilibrium is static.
- Eulerian (stationary) equilibria can have flow. Their stability can be studied with "Eulerianized" Lagrangian displacements (ELD).
- Another method involves Dynamically Accessible Variations (DAV), which are constrained to satisfy the invariants of the flow. Closely related to the Energy–Casimir method.
- We show the equivalence of the the ELD and DAV methods for the case of MHD equilibria.

History: Some key papers

- Fjørtoft [1950] Geophysical context.
- Lundquist [1951], Bernstein et al. [1958], Woltjer [1958] Static equilibria.
- Kruskal and Oberman [1958], Gardner [1963] Kinetic theory.
- Arnold [1965, 1966, 1969] Nonlinear stability criterion.
- Frieman and Rotenberg [1960], Newcomb [1962], Hameiri [1982, 1998], Hameiri and Holties [1994] Stationary equilibria, Nonlinear eigenvalue problem.
- Holm et al. [1985], Morrison and Eliezer [1986], Finn and Sun [1987] Energy–Casimir method).
- Morrison and Pfirsch [1990], Morrison [1998] Dynamical accessibility.

Equations of Motion

Inviscid, ideally conducting fluid:

$$\rho \left(\partial_t \boldsymbol{v} + \boldsymbol{v} \cdot \nabla \boldsymbol{v}\right) = -\nabla p + \boldsymbol{j} \times \boldsymbol{B},$$
$$\partial_t \rho + \nabla \cdot (\rho \, \boldsymbol{v}) = 0,$$
$$\partial_t s + \boldsymbol{v} \cdot \nabla s = 0,$$
$$\partial_t \boldsymbol{B} - \nabla \times (\boldsymbol{v} \times \boldsymbol{B}) = 0.$$

Conserved energy (Hamiltonian):

$$H = \int d^3x \left(\frac{1}{2} \rho |\mathbf{v}|^2 + \frac{1}{2} |\mathbf{B}|^2 + \rho U(\rho, s) \right)$$

Can have other invariants, such as the helicity and cross-helicity, depending on initial configuration (Padhye and Morrison [1996], Hameiri [1998]).

Equilibrium quantities are denoted by a subscript, "e". Setting ∂_t and v_e to zero, the only condition is

$$\nabla p_{\rm e} = (\nabla \times \boldsymbol{B}_{\rm e}) \times \boldsymbol{B}_{\rm e}, \qquad \nabla \cdot \boldsymbol{B}_{\rm e} = 0.$$

To determine a sufficient condition for stability, we consider perturbations about a static equilibrium

$$\boldsymbol{x} = \boldsymbol{x}_0 + \boldsymbol{\xi}(\boldsymbol{x}_0, t),$$

where x is the position of a fluid element at time t and $\xi(x_0, t)$ is the Lagrangian displacement, with $\xi(x_0, 0) = 0$.

After computing the variations of the various physical quantities and linearizing the equations of motion with respect to $\boldsymbol{\xi}$ (Bernstein et al. [1958]), we obtain

 $\rho_0 \, \ddot{\boldsymbol{\xi}} = \mathbf{F}(\boldsymbol{\xi}).$

(Formal) linear stability is then guaranteed if

$$\delta W(\boldsymbol{\xi}, \boldsymbol{\xi}) \coloneqq -\frac{1}{2} \int \boldsymbol{\xi} \cdot \mathbf{F}(\boldsymbol{\xi}) \, \mathrm{d}^3 x \ge 0.$$

This is Lagrange's principle: the potential energy needs to be positive-definite for stability.

The relabeling symmetry allows passage from the Lagrangian to the Eulerian picture via the process of reduction (Marsden and Weinstein [1974], Morrison [1998]). The equilibria then represent stationary flows. Three approaches:

- "Eulerianized" Lagrangian displacements (Frieman and Rotenberg [1960], Newcomb [1962]), by which the displacements are re-expressed in terms of Eulerian variables only.
- Energy–Casimir Method (Holm et al. [1985], Morrison and Eliezer [1986]).
- Dynamically accessible variations (Morrison and Pfirsch [1990], Morrison [1998]), a method for generating variations which preserve the Casimir invariants of the system

Express the Lagrangian displacement $\boldsymbol{\xi}(\boldsymbol{x}_0, t)$ in terms of the Eulerian coordinates \boldsymbol{x} :

$$\boldsymbol{\eta}(\boldsymbol{x},t) = \boldsymbol{\xi}(\boldsymbol{x}_0,t)$$
 (back)

The variations are (Newcomb [1962])

$$egin{aligned} \delta oldsymbol{v} &= \dot{oldsymbol{\eta}} + oldsymbol{v} \cdot
abla oldsymbol{\eta} - oldsymbol{\eta} \cdot
abla oldsymbol{v}, \ \delta oldsymbol{
ho} &= - oldsymbol{
abla} \cdot (oldsymbol{
ho} oldsymbol{\eta}), \ \delta oldsymbol{s} &= -oldsymbol{\eta} \cdot
abla oldsymbol{s}, \ \delta oldsymbol{B} &=
abla imes (oldsymbol{\eta} imes oldsymbol{B}). \end{aligned}$$

Energy can be varied with respect to these perturbations: a sufficient stability criterion is obtained. η and $\dot{\eta}$ are independent.

Ideal MHD has a Hamiltonian formulation in terms of a noncanonical bracket (Morrison and Greene [1980])

$$\{F, G\} = -\left(\int d^3x \, F_\rho \, \nabla \cdot \, G_{\boldsymbol{v}} + F_{\boldsymbol{v}} \cdot \left(\frac{(\nabla \times \boldsymbol{v})}{2\rho} \times G_{\boldsymbol{v}}\right) + \rho^{-1} \, \nabla s \cdot (F_s \, G_{\boldsymbol{v}}) + \rho^{-1} F_{\boldsymbol{v}} \cdot (\boldsymbol{B} \times (\nabla \times G_{\boldsymbol{B}}))\right) + \left(F \longleftrightarrow G\right)$$

F and G are functionals of the dynamical variables (v, ρ, s, B) , and subscripts denote functional derivatives. The bracket $\{,\}$ is antisymmetric and satisfies the Jacobi identity. The equations of motion can be written

$$\partial_t(\boldsymbol{v},\rho,s,\boldsymbol{B}) = \{(\boldsymbol{v},\rho,s,\boldsymbol{B}),H\}.$$

Another method establishing formal stability uses dynamically accessible variations (DAV), defined for the variable ζ as

$$\delta\zeta_{\mathrm{da}} \coloneqq \{\mathcal{G}, \zeta\}, \qquad \delta^2\zeta_{\mathrm{da}} \coloneqq \frac{1}{2} \{\mathcal{G}, \{\mathcal{G}, \zeta\}\},$$

with \mathcal{G} given in terms of the generating functions χ_{μ} by

$$\mathcal{G} \coloneqq \sum_{\mu} \int \zeta^{\mu} \, \chi_{\mu} \, \mathrm{d}^{3} x.$$

DAV are variations that are constrained to remain on the symplectic leaves of the system. They preserve the Casimir invariants to second order (but there is no need to explicitly know the invariants).

Stationary solutions ζ_e of the Hamiltonian,

$$\delta H_{\rm da}[\zeta_{\rm e}] = 0,$$

capture all possible equilibria of the equations of motion. The energy of the perturbations is

$$\delta^2 H_{\rm da}[\zeta_{\rm e}] = \frac{1}{2} \int \left(\delta \zeta_{\rm da}^{\sigma} \frac{\delta^2 H}{\delta \zeta^{\sigma} \,\delta \zeta^{\tau}} \,\delta \zeta_{\rm da}^{\tau} + \delta^2 \zeta_{\rm da}^{\nu} \frac{\delta H}{\delta \zeta^{\nu}} \right) \,\mathrm{d}^3 x,$$

with $\zeta = (\boldsymbol{v}, \rho, s, \boldsymbol{B})$ and repeated indices are summed. Positive-definiteness of $\delta^2 H_{da}[\zeta_e]$ implies formal stability, which implies linear stability, but not nonlinear stability. (Requires convexity, Holm et al. [1985].)

DAV for MHD

The form of the dynamically accessible variations is

$$\begin{split} \rho \, \delta \boldsymbol{v}_{\mathrm{da}} &= (\nabla \times \boldsymbol{v}) \times \boldsymbol{\chi}_0 + \rho \, \nabla \chi_1 - \chi_2 \, \nabla s + \boldsymbol{B} \times (\nabla \times \boldsymbol{\chi}_3), \\ \delta \rho_{\mathrm{da}} &= \nabla \cdot \boldsymbol{\chi}_0, \\ \delta s_{\mathrm{da}} &= \rho^{-1} \, \boldsymbol{\chi}_0 \cdot \nabla s, \\ \delta \boldsymbol{B}_{\mathrm{da}} &= \nabla \times \left(\rho^{-1} \, \boldsymbol{B} \times \boldsymbol{\chi}_0 \right). \end{split}$$

 χ_0, χ_1, χ_2 , and χ_3 are the arbitrary generating functions of the variations. The variations for ρ , s, and B are the same as for the ELD, with $\chi_0 = -\rho \eta$.

The combination of arbitrary functions in the definition of δv_{da} makes that perturbation arbitrary, in the same manner as the ELD perturbation δv , as we now show.

The compelling choice is $\chi_0 = \rho \eta$, from which the equivalence of the *v* perturbations requires that

$$\dot{\boldsymbol{\eta}} = \rho \, \nabla \chi_1 - \chi_2 \, \nabla s + \boldsymbol{B} \times (\nabla \times \boldsymbol{\chi}_3).$$

The ELD and the DAV will be equivalent if it is possible to choose χ_1 , χ_2 , and χ_3 to span the same space as $\dot{\eta}$, and vice-versa.

 $\dot{\eta}$ can represent any perturbation, up to boundary conditions.

Local Euler–Clebsch representation for magnetic field:

$$\boldsymbol{B} = \nabla \boldsymbol{\alpha} \times \nabla \boldsymbol{\beta} \quad [+\nabla \boldsymbol{\gamma} \times \nabla \Psi(\boldsymbol{\alpha}, \boldsymbol{\beta}, \boldsymbol{\gamma})]$$

[More generally, Boozerize.]

Pick a third, independent function γ . Covariant representation:

$$\chi_{3} = a \,\nabla\alpha + b \,\nabla\beta + c \,\nabla\gamma$$
$$\nabla \times \chi_{3} = \nabla a \times \nabla \alpha + \nabla b \times \nabla \beta + \nabla c \times \nabla\gamma$$
$$\boldsymbol{B} \times (\nabla \times \chi_{3}) = J \left(\frac{\partial a}{\partial \gamma} - \frac{\partial c}{\partial \alpha}\right) \nabla \alpha - J \left(\frac{\partial c}{\partial \beta} - \frac{\partial b}{\partial \gamma}\right) \nabla \beta$$
$$J \coloneqq \nabla \alpha \cdot (\nabla \beta \times \nabla \gamma)$$
$$\dot{\boldsymbol{\eta}} = \left(\rho \frac{\partial \chi_{1}}{\partial \alpha} - \chi_{2} \frac{\partial s}{\partial \alpha} + J \left(\frac{\partial a}{\partial \gamma} - \frac{\partial c}{\partial \alpha}\right)\right) \nabla \alpha$$
$$- \left(\rho \frac{\partial \chi_{1}}{\partial \beta} - \chi_{2} \frac{\partial s}{\partial \beta} - J \left(\frac{\partial c}{\partial \beta} - \frac{\partial b}{\partial \gamma}\right)\right) \nabla \beta + \left(\rho \frac{\partial \chi_{1}}{\partial \gamma} - \chi_{2} \frac{\partial s}{\partial \gamma}\right) \nabla$$

Covariant representation of $\dot{\eta}$:

$$\dot{\boldsymbol{\eta}} = A \, \nabla \alpha + B \, \nabla \beta + C \, \nabla \gamma$$

Equate coefficients:

$$\rho \frac{\partial \chi_1}{\partial \alpha} - \chi_2 \frac{\partial s}{\partial \alpha} + J \left(\frac{\partial a}{\partial \gamma} - \frac{\partial c}{\partial \alpha} \right) = A \tag{A}$$

$$\rho \frac{\partial \chi_1}{\partial \beta} - \chi_2 \frac{\partial s}{\partial \beta} - J \left(\frac{\partial c}{\partial \beta} - \frac{\partial b}{\partial \gamma} \right) = B \tag{B}$$

$$\rho \frac{\partial \chi_1}{\partial \gamma} - \chi_2 \frac{\partial s}{\partial \gamma} = C \tag{C}$$

The function a only appears in (A), so solve for $\partial a/\partial \gamma$ and integrate; b only appears in (B), so solve for $\partial b/\partial \gamma$ and integrate. Use χ_2 to satisfy (C).

Conclusions

- The three approaches, using Lagrangian perturbations vs energy–Casimir and dynamical accessibility, lead to essentially the same stability criterion.
- The dynamical accessibility method can be used directly at the Hamiltonian level. One needs to know the Poisson bracket and Hamiltonian.
- For energy–Casimir, one also needs the Casimir invariants, but not necessarily the bracket.
- In both approaches other invariants (non-Casimir, e.g., momentum) can be incorporated.
- Dynamical accessibility has also been applied to Vlasov–Maxwell equilibria (Morrison and Pfirsch [1989, 1990]).

References

- V. I. Arnold. Conditions for nonlinear stability of the stationary plane curvilinear flows of an ideal fluid. *Doklady Mat. Nauk.*, 162(5):773–777, 1965.
- V. I. Arnold. Sur un principe variationnel pour les écoulements stationaires des liquides parfaits et ses applications aux problèmes de stabilité non linéaires. *Journal de Mécanique*, 5:29–43, 1966.
- V. I. Arnold. On a priori estimate in the theory of hydrodynamic stability. Am. Math. Soc. Transl., 19:267–269, 1969.
- I. B. Bernstein, E. A. Frieman, M. D. Kruskal, and R. M. Kulsrud. An energy principle for hydromagnetic stability problems. *Proc. R. Soc. Lond. A*, 244:17–40, 1958.
- J. M. Finn and Guo-Zheng Sun. Nonlinear stability and the energy-Casimir method. *Comments Plasma Phys. Controlled Fusion*, 11(1):7–25, 1987.
- Ragnar Fjørtoft. Application of integral theorems in deriving criteria for stability for laminar flows and for the baroclinic circular vortex. *Geofys. Pub.*, 17(6):1–52, 1950.
- E. A. Frieman and M. Rotenberg. On hydrodynamic stability of stationary equilibria. *Rev. Mod. Phys.*, 32(4):898–902, October 1960.

Clifford S. Gardner. Bound on the energy available from a plasma. Phys. Fluids, 6(6):839-840, June 1963.

- E. Hameiri. The equilibrium and stability of rotating plasmas. Phys. Fluids, 26(1):230-237, January 1982.
- E. Hameiri. Variational principles for equilibrium states with plasma flow. Phys. Plasmas, 5(9):3270-3281, September 1998.
- E. Hameiri and H. A. Holties. Improved stability conditions for rotating plasmas. *Phys. Plasmas*, 1(12):3807–3813, December 1994.
- Darryl D. Holm, Jerrold E. Marsden, Tudor Ratiu, and Alan Weinstein. Nonlinear stability of fluid and plasma equilibria. *Physics Reports*, 123(1 & 2):1–116, July 1985.
- M. D. Kruskal and C. R. Oberman. On the stability of a plasma in static equilibrium. Phys. Fluids, 1:275-280, 1958.
- S. Lundquist. On the stability of magneto-hydrostatic fields. Phys. Rev., 83(2):307-311, July 1951.
- Jerrold E. Marsden and Alan Weinstein. Reduction of symplectic manifolds with symmetry. *Rep. Math. Phys.*, 5(1):121–130, 1974.

Philip J. Morrison and S. Eliezer. Spontaneous symmetry breaking and neutral stability in the noncanonical Hamiltonian formalism. *Phys. Rev. A*, 33(6):4205–4214, June 1986.

Philip J. Morrison. Hamiltonian description of the ideal fluid. Rev. Mod. Phys., 70(2):467–521, April 1998.

- Philip J. Morrison and John M. Greene. Noncanonical Hamiltonian density formulation of hydrodynamics and ideal magnetohydrodynamics. *Phys. Rev. Lett.*, 45(10):790–794, September 1980.
- Philip J. Morrison and D. Pfirsch. Free-energy expressions for Vlasov equilibria. *Phys. Rev. A*, 49(7):3898–3910, October 1989.

Philip J. Morrison and D. Pfirsch. The free energy of Maxwell-Vlasov equilibria. Phys. Fluids B, 2(6):1105–1113, June 1990.

W. A. Newcomb. Lagrangian and Hamiltonian methods in magnetohydrodynamics. *Nuclear Fusion: Supplement, part 2*, pages 451–463, 1962.

Nikhil Padhye and Philip J. Morrison. Fluid element relabeling symmetry. Phys. Lett. A, 219:287–292, 1996.

L. Woltjer. The stability of force-free magnetic fields. Astrophys. J., 128:384–391, 1958.

Expression for $F(\boldsymbol{\xi})$

 $\boldsymbol{Q}\coloneqq
abla_0 imes (\boldsymbol{\xi} imes \boldsymbol{B}_0)$

(back)