

# Mixing in Thin Flows over a Curved Substrate

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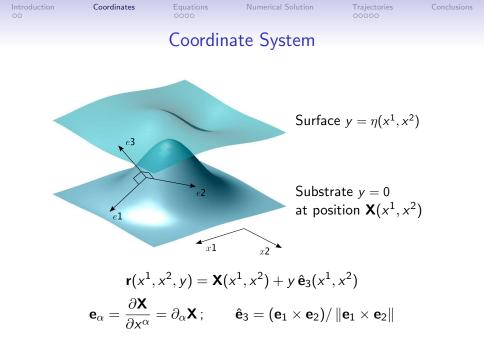
APS-DFD, Tampa Bay, 20 November 2006



- A thin layer of fluid flowing down an inclined substrate.
- Reduce to two-dimensional problem by asymptotic expansion: PDE for the height field.
- But the velocity field is still three-dimensional, with a nontrivial vertical component.
- Steady three-dimensional flows can exhibit chaotic trajectories.
- This leads to fluid particles rapidly decorrelating: good for mixing.
- Can suitable substrate shapes lead to good horizontal mixing?



- Thin-layer expansion in the direction normal to the substrate.
- Similar derivation to [Roy, Roberts, and Simpson, JFM 454, 235 (2002)].
- For simplicity, assume steady flow.
- Use non-orthogonal coordinates, since globally orthogonal coordinates are difficult to compute in general.
- Correct velocity field to satisfy kinematic constraints this is crucial for particle advection.
- Integrate trajectories and make Poincaré sections in a spatially periodic domain.



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## **Dynamical Equations**

We now assume  $\boldsymbol{u}$  satisfies the Stokes equation,

$$\Delta \mathbf{u} = \nabla p - \hat{\mathbf{g}}, \qquad \nabla \cdot \mathbf{u} = \mathbf{0},$$

where p is the pressure and  $\hat{\mathbf{g}}$  is a unit vector in the direction of gravity. The velocity satisfies the boundary conditions

where

$$au \coloneqq 
abla \mathbf{u} + (
abla \mathbf{u})^T$$

is the deviatoric stress,  $\hat{\mathbf{n}}$  is the unit normal to the surface,  $\mathbf{t}_{\alpha}$  are tangents to the surface, and  $\kappa_{\mathrm{surf}}$  is the mean curvature of the surface. All quantities are dimensionless.



The time has come to make the layer thin: we do this by assuming that horizontal scales vary slowly:

$$x^{\alpha} = \varepsilon^{-1} x^{\alpha *}, \quad v = \varepsilon v^{*}, \quad p = \varepsilon^{-1} p^{*}, \quad \sigma = \varepsilon^{-2} \sigma^{*}.$$

Everything else is of order unity, including vertical scales. We immediately drop the \* superscripts, and expand the fields as

$$u^{\alpha} = u^{\alpha}_{(0)} + \varepsilon u^{\alpha}_{(1)} + \dots,$$
  
$$p = p_{(0)} + \varepsilon p_{(1)} + \dots$$

Note that we leave v unexpanded (see later).

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#### The Mass Flux

The horizontal velocity field is sufficient to find the mass flux,

$$\begin{split} q^{\alpha} &= \int_{0}^{\eta} \omega \, u^{\alpha} dy = \int_{0}^{\eta} (1 - \varepsilon \, \kappa \, y) u^{\alpha} \, dy + \mathcal{O}(\varepsilon^{2}) \,, \\ &= q^{\alpha}_{\text{grav}} + q^{\alpha}_{\text{surf}} \,, \end{split}$$

$$\begin{split} q_{\rm grav}^{\alpha} &= \frac{1}{3} \eta^3 \left\{ \hat{g}_{\rm s}^{\alpha} - \varepsilon \, \hat{g}_{\rm s}^{\beta} \left( \kappa \, \delta_{\beta}{}^{\alpha} + \frac{1}{2} \, \mathbb{K}_{\beta}{}^{\alpha} \right) \eta + \varepsilon \, \hat{g}_{y} \, \partial^{\alpha} \eta \right\} \\ &+ \varepsilon^2 \frac{1}{120} \, \eta^4 \kappa \left\{ \eta \, \hat{g}_{\rm s}^{\beta} \left( 9 \kappa \, \delta_{\beta}{}^{\alpha} + 11 \, \mathbb{K}_{\beta}{}^{\alpha} \right) - 25 \, \hat{g}_{y} \, \partial^{\alpha} \eta \right\} + \mathcal{O} \left( \varepsilon^2 \right), \end{split}$$

$$\begin{split} q^{\alpha}_{\rm surf} &= \frac{1}{3} \sigma \eta^3 \left\{ \partial^{\alpha} \kappa_{\rm surf} - \varepsilon \, \eta \, \kappa \, \partial^{\alpha} \kappa + \frac{1}{2} \varepsilon \, \eta \, \mathbb{K}_{\beta}{}^{\alpha} \, \partial^{\beta} \kappa \right\} \\ &+ \varepsilon^2 \frac{1}{120} \, \sigma \, \eta^4 \kappa \left\{ 9 \eta \, \kappa \, \partial^{\alpha} \kappa - 14 \, \eta \mathbb{K}_{\beta}{}^{\alpha} \partial^{\beta} \kappa - 25 \, \partial^{\alpha} (\kappa_2 \eta + \Delta \eta) \right\} + \mathcal{O} \big( \varepsilon^2 \big), \end{split}$$

Note that we've kept some second-order terms but not others. The above fluxes are only asymptotic to order  $\varepsilon^1$ , but they preserve the free-surface kinematic BC to all orders...

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### The Vertical Velocity

The vertical velocity is obtained from mass conservation:

$$v = -rac{1}{1-arepsilon\kappa y}\int_0^y \partial_lpha \left( \left(1-arepsilon\kappa y
ight) u^lpha 
ight) \, dy \,, \quad {
m not} \; {
m expanded} \; {
m in} \; arepsilon.$$

Mass conservation follows from using this form for v, and the free-surface kinematic boundary condition is satisfied exactly if the second-order terms are included in the flux.

The exact kinematic constraints are crucial for particle advection:

- Mass preservation prevents the existence of attractors in the flow where particles bunch up.
- The kinematic boundary condition prevents particles escaping from the top surface of the flow.

These are only exact to the extent that  $\nabla_{\alpha}q^{\alpha} = 0$  is satisfied numerically, but this is a much smaller error than  $\varepsilon^2$ .

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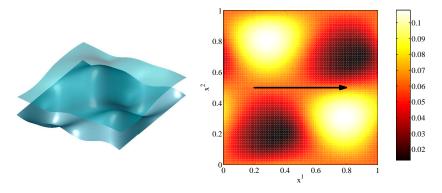
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#### Numerical Solution

We now solve  $\nabla_{\alpha}q^{\alpha} = 0$  for the height field  $\eta(x^1, x^2)$ . The pictures below are for a substrate shape given by

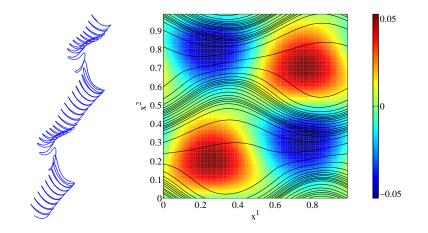
 $f(x^{1}, x^{2}) = 0.05 \{ \sin(2\pi x^{1}) \sin(2\pi x^{2}) + 0.2 \sin(4\pi x^{2}) \}$ 



Parameters:  $\varepsilon = 0.06$  (thickness),  $\theta = 0.1$  (tilt),  $\phi = 0$ ,  $\sigma = 0$ .



#### A Typical Regular Trajectory



- The particle explores the top and bottom of the layer.
- Nonchaotic, but regularly 'jumps' laterally in  $x^2$ .

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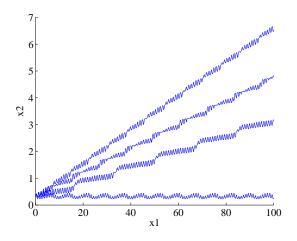
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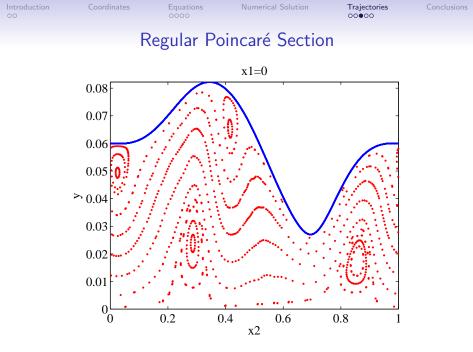
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#### Two Types of Trajectories

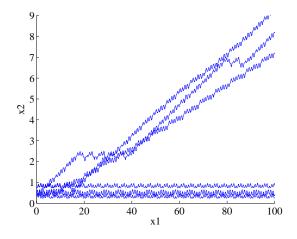


- Trapped and untrapped trajectories.
- The effective diffusivity is strongly anisotropic.

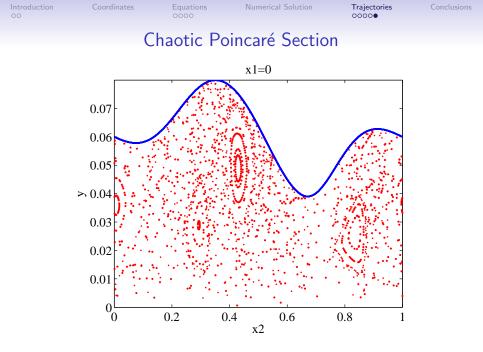


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Decrasing the Tilt Angle: Chaotic Trajectories



- The trajectories chaotically 'jump' between channels.
- Sequence of Lévy flights and trapped orbits, analogous to Solomon, Weeks, & Swinney (1993).





### Conclusions

- Wide range of behaviour of particle trajectories.
- This will affect mixing properties: strongly anisotropic effective diffusivity.
- Important for coating applications?
- Since the flow is 3D, can have chaos. Three factors:
  - 1. Break discrete symmetries of the substrate;
  - 2. Make the fluid deeper (but still thin);
  - 3. Make the tilt angle  $\theta$  smaller.
- Relate substrate properties (curvature tensor) to chaotic features.
- Experiments!
- Time-dependence: induce chaotic mixing by vibrating the substrate or sending waves through it.