

Cloud–Radiation Feedback and Atmosphere–Ocean Coupling in a Stochastic Multicloud Model

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Abstract

Despite recent advances in supercomputing, current general circulation models (GCMs) have significant problems in representing the variability associated with organized tropical convection. Furthermore, due to high sensitivity of the simulations to the cloud radiation feedback, the tropical convection remains a major source of uncertainty in long-term weather and climate forecasts. In a series of recent studies, it has been shown, in paradigm two-baroclinic-mode systems and in aquaplanet GCMs, that a stochastic multicloud convective parameterization based on three cloud types (congestus, deep and stratiform) can be used to improve the variability and the dynamical structure of tropical convection, including intermittent coherent structures such as synoptic and mesoscale convective

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systems. Here, the stochastic multcloud model is modified with a parameterized cloud radiation feedback mechanism and atmosphere-ocean coupling. The radiative convective feedback mechanism is shown to increase the mean and variability of the Walker circulation. The corresponding intensification of the circulation is associated with propagating synoptic scale systems originating inside of the enhanced sea surface temperature area. In column simulations, the atmosphere ocean coupling introduces pronounced low frequency convective features on the time scale associated with the depth of the mixed ocean layer. However, in the presence of the gravity wave mixing of spatially extended simulations, these features are not as prominent. This highlights the deficiency of the column model approach at predicting the behavior of multiscale spatially extended systems. Overall, the study develops a systematic framework for incorporating parameterized radiative cloud feedback and ocean coupling which may be used to improve representation of intraseasonal and seasonal variability in GCMs.

Keywords: Stochastic convective parameterization, multcloud models, tropical atmospheric dynamics, convectively coupled waves, cloud radiation feedback, atmosphere ocean coupling

1. Introduction

Atmospheric dynamics in the tropics are characterized by the predominance of organized convection on a wide range of scales, spanning mesoscale systems to synoptic and planetary-scale convectively coupled waves such as Kelvin waves
5 and the Madden Julian oscillation (MJO) (Nakazawa, 1974; Hendon and Liebmann, 1994; Wheeler and Kiladis, 1999). While the importance of the tropics to weather and climate forecasts cannot be overestimated, present coarse resolution GCMs used for the prediction of weather and climate have significant problems in representing variability associated with tropical convection (Slingo
10 et al., 1996; Moncrieff and Klinker, 1997; Scinocca and McFarlane, 2004; Lau and Waliser, 2005; Zhang, 2005).

It is believed that the deficiency is due to insufficient treatment of the cumulus convection (Moncrieff and Klinker, 1997; Lin et al., 2006), which has to be parameterized in GCMs. The inaccuracy of the vertical and horizontal cloud
15 distributions furthers the already great uncertainty associated with cloud radiation feedback (CRF) mechanisms (Stephens and Webster, 1979; Cess et al., 1990, 1996; Bony and Emanuel, 2005) and the large-scale atmospheric circulation (Peters and Bretherton, 2005; Tian and Ramanathan, 2003). Given the sensitivity of tropical variability (Stephens and Webster, 1979; Cess et al., 1990,
20 1996; Bony and Emanuel, 2005) to CRF, the search for new strategies for the parameterization of tropical convection and associated radiative feedback and atmosphere ocean coupling (AOC) effects is one of the central problems in the atmospheric community.

Several methods have been developed to address the multiscale nature of
25 tropical convection. Cloud-resolving models (CRM) on fine computational grids have succeeded in representing some aspects of organized convection, including mesoscale organization and cloud distribution (ECMWF, 2003; Moncrieff et al., 2007; Slawinska et al., 2014b). In addition, superparameterization (SP) methods (Grabowski and Smolarkiewicz, 1999; Grabowski, 2001, 2004; Randall
30 et al., 2003; Majda, 2007) and sparse space-time SP (Xing et al., 2009; Slawinska et al., 2014a) use a cloud resolving model in each column of the large scale GCM. In the SP setup, the CRM is used to explicitly represent small scale processes such as cloud microphysics and convective updrafts, thereby representing an extended range of the scales of tropical convection compared
35 to a traditional GCM. However, despite considerable success in duplicating observed radiative fluxes (Wu and Moncrieff, 2000), these methods are not currently computationally viable for application to climate simulations. Moreover, these complex models do not necessarily further the qualitative understanding of the processes involved. For example, CRF is usually computed through highly complex multiple-scattering and radiative transfer models (Stephens and Webster,
40 1979; Zurovac-Jevtik et al., 2005).

Another novel approach to the problem of missing tropical variability in

GCMs has been the development of the multicloud parameterizations (Khouider and Majda, 2006b,a, 2007, 2008a,b; Khouider et al., 2010; Frenkel et al., 2012; Peters et al., 2013), which capture the interaction of the three cloud types (congestus, deep and stratiform) which characterize tropical convection. In particular, the stochastic multicloud model (Khouider et al., 2010; Frenkel et al., 2012, 2013) (hereafter KBM10, FMK12, FMK13) aims to capture these phenomena with a Markov chain lattice model where each lattice site is either occupied by a cloud of a certain type or is considered a clear-sky site. The convective elements interact with the large-scale environment and with each other through convective available potential energy (CAPE) and middle troposphere dryness. When local interactions between the individual lattice sites are ignored, or when nearest-neighbor interactions are allowed (Khouider, 2014), the dynamical evolution of the cloud area fractions in the stochastic multicloud model takes the form of a computationally inexpensive coarse grained stochastic process (Katsoulakis et al., 2003; Khouider et al., 2003; Majda et al., 2008). In addition to an enhanced representation of clouds, the framework is simple enough to allow semi-analytic solutions. In particular these authors were able to study stability and bifurcations of the solutions attributable to the diurnal surface fluxes (Frenkel et al., 2010). Despite its apparent simplicity, the multicloud model is very successful in capturing most of the Wheeler-Kiladis-Takayabu spectrum of convectively coupled waves (Takayabu, 1994; Wheeler and Kiladis, 1999) in terms of linear wave theory (KM06a, KM08a, Han and Khouider (2010)) and nonlinear organization of large-scale envelopes mimicking cross-scale interactions of the Madden-Julian oscillation (MJO) and convectively coupled waves (KM07; KM08b; Majda et al. (2007)), in the idealized context of a simple two-baroclinic modes model employed here. Both the deterministic and stochastic multicloud models dramatically improve the representation of the coherent and intermittent nature of organized convection. This has been shown in an idealized two baroclinic mode framework coupled to a coarse resolution GCM for both the MJO and monsoon intraseasonal oscillations (Khouider et al., 2011; Deng et al., 2014; Ajayamohan et al., 2013, 2014; Deng et al., 2014).

Here, a version of the stochastic multcloud parameterization (FMK13) is
75 augmented with cloud radiation feedback and ocean coupling mechanism. It is
natural to use the stochastic, rather than deterministic, multcloud parameter-
ization for this purpose; this is because cloud fractions play such an important
role in cloud–radiation feedback, and cloud fractions were introduced into the
multicloud model as part of the stochastic parameterization. As in FMK13, the
80 parameterization is coupled to a simplified model of the primitive equations; the
vertical resolution is reduced to the first two baroclinic modes. Such a setup
is intermediate between overly simple one baroclinic mode models (Tian and
Ramanathan, 2003; Sobel et al., 2004) and more complex GCM and CRM sim-
ulations (Zurovac–Jevtik et al., 2005; Moncrieff and Klinker, 1997). The impact
85 of radiative convective feedback of each of the three cloud types is parameterized
through a product of the cloud fraction and two parameters, which represent the
idealized projection of the cloud radiative feedback effect onto the two baroclinic
modes of the system. The radiative feedback is shown to increase the strength
and variability of the Walker circulation. The atmosphere-ocean coupling in-
90 creases the variability of convection by introducing low frequency envelopes of
synoptic and mesoscale convective systems. Single column simulations are used
here to isolate and elucidate the effects of these modifications. Atmosphere-
ocean coupling and cloud radiation feedback have a subtle but significant effect
in spatially extended simulations. Spectral analysis highlights the effects of the
95 modifications introduced here and their interactions with intrinsic variability of
the system.

The remainder of the paper is organized as follows. A self-contained review
of the stochastic multcloud parameterization is presented in Section 2. The
section also introduces cloud radiation feedback and atmosphere ocean coupling
100 mechanisms. In Section 3, single column simulations are used to illustrate the
effects of the two modifications listed above. In Section 4, the modified param-
eterization is used to study flows above the equator in a series of idealized Walker
cell simulations. Some discussion and concluding remarks are given in Section
5.

105 2. Multicloud model, cloud radiation feedback and atmosphere-ocean coupling

We start with a brief review of the dynamical core equations used for the stochastic multicloud parameterization in Section 2.1. A more thorough and detailed discussion of the model equations is found in (KM06). Nevertheless,
 110 a comprehensive list of the model constants and parameters is given in Table 1 for the sake of completeness. Section 2.2 reviews the stochastic multicloud parameterization, while Sections 2.3 and 2.4 introduce cloud radiative feedback and ocean coupling, respectively.

2.1. Dynamical core

115 The stochastic multicloud parameterization assumes three heating profiles associated with the main cloud types that characterize organized tropical convective systems (Johnson et al., 1999): cumulus congestus clouds that heat the lower troposphere and cool the upper troposphere, through radiation and de-
 trainment, deep convective towers that heat the whole tropospheric depth, and
 120 the associated lagging-stratiform anvils heat the upper troposphere and cool the lower troposphere, due to evaporation of stratiform rain. Accordingly, the dynamical core used in this paper consists of two coupled and forced shallow water systems. Without the meridional dependency, the equations are given by

$$\partial_t u_j - \partial_x \theta_j = C_d u_0 u_j - \frac{1}{\tau_R} u_j, \quad j = 1, 2 \quad (1)$$

$$\partial_t \theta_1 - \partial_x u_1 = H_d + \xi_s H_s + \xi_c H_c + S_1, \quad (2)$$

$$125 \quad \partial_t \theta_2 - \frac{1}{4} \partial_x u_2 = H_c - H_s + S_2. \quad (3)$$

Here H_d , H_s and H_c are the heating rates for deep, stratiform and cumulus congestus clouds obtained by either the deterministic or the stochastic parameterization. These heating rates are combined to form the bulk precipitation $P = H_d + \xi_s H_s + \xi_c H_c$. The coefficients ξ_c and ξ_s denote contribution of con-
 gestus and stratiform rain to the bulk precipitation. The parameters C_d and u_0
 130 are respectively the momentum drag coefficient and the strength of turbulent

fluctuations in the boundary layer. The last terms of the first and second baroclinic heating mode equations, $S_j, j = 1, 2$, represent radiative source terms, including cloud–radiation feedback, and will be discussed in Section 2.3.

135 The multicloud models additionally carry an equation for the vertically integrated tropospheric moisture content, q , and an equation for the boundary layer equivalent potential temperature, θ_{eb} .

$$\partial_t q + \partial_x[(u_1 + \tilde{\alpha}u_2)q + (u_1 + \tilde{\lambda}u_2)\tilde{Q}] = -\frac{2\sqrt{2}}{\pi}P + D/H_T \quad (4)$$

$$\partial_t \theta_{eb} = \frac{1}{h_b}(E - D). \quad (5)$$

The sea surface saturation equivalent potential temperature, $\theta_{eb}^*(T_s)$, is a function of the bulk ocean layer temperature discussed in Section 2.4 . For simulations without AOC, $\theta_{eb}^*(T_s)$ is set to constant ,so that $\bar{\theta}_{eb}^* - \bar{\theta}_{eb} = 10K$. Here and throughout the paper, \bar{X} denotes the radiative–convective equilibrium (RCE) value of the variable X . Notice that the simple treatment (5) of the boundary layer does not explicitly include the mechanism of Lindzen and Nigam (1987). In
140 the (x, t) simulations where SST is not homogenous, the sea surface evaporation E takes the form

$$\frac{E}{h_b} = \tau_e^{-1}(\theta_{eb}^*(T_s) + \theta_{eb\Delta}^*(x) - \theta_{eb}). \quad (6)$$

The sea surface saturation equivalent potential temperature takes the form

$$\theta_{eb\Delta}^*(x) = 5 \cos\left(\frac{4\pi x}{40000}\right) + 10K, \quad (7)$$

within an interval of 20,000 km of the 40,000 km domain and $\theta_{eb}^* = 5$ K everywhere else as in Khouider and Majda (2007) and KM08a. This setup mimics
150 the Indian Ocean–Western Pacific warm pool.

2.2. Stochastic multicloud parameterization

The stochastic multicloud parameterization is designed to capture the dynamical interactions between the three cloud types that characterize organized tropical convection and the environment. In the stochastic multicloud model
155 these interactions are represented through a coarse grained lattice model (KBM10).

Parameter	Value	Description
$h_b/H_m/H_T$	500 m / 5 km/ 16 km	depth of ABL/ mid-troposphere/ free troposphere
ξ_s/ξ_c	0.4/0	Stratiform/Congestus contribution to first baroclinic mode
\tilde{Q}	0.9	Background moisture stratification
$\tilde{\lambda}/\tilde{\alpha}$	0.8/0.1	Coefficient of u_2 in linear / nonlinear moisture convergence
m_0	Determined at RCE	Large-scale background downdraft velocity scale
μ	0.25	Contribution of convective downdrafts to D
α_s/α_c	0.25/ 0.1	Stratiform/Congestus adjustment coefficient
τ_R/τ_D	75 days / 50 days	Rayleigh drag / Newtonian cooling time scale
τ_s/τ_c	3 hours / 2 hour	Stratiform /Congestus adjustment time scale
τ_{conv}	2 hours	Convective time scale
τ_e	Determined by RCE	Surface evaporation time scale
\bar{Q}	Determined at RCE	Bulk convective heating at RCE
$\bar{\theta}_{eb} - \bar{\theta}_{em}$	11 K	Mean (RCE) Dryness of the atmosphere
a_1/a_2	0.45 / 0.55	Relative contribution of θ_{eb} / q to deep convection
a_0/a'_0	2 / 1.5	Dry convective buoyancy frequency in deep/congestus eqns.
γ_2/γ'_2	0.1 / 2	Relative contribution of θ_2 to deep /congestus heating
α_2	0.1	Relative contribution of θ_2 to θ_{em}
C_d	0.001	Surface drag coefficient
u_0	2 m/s	Strength of turbulent fluctuations
$CAPE_0$	400 J/Kg	Reference values of CAPE
T_0	12 K	Reference values of dryness
$\bar{\alpha}$	≈ 15 K	Unit scale of temperature
R_{univ}	≈ 8.31436 J/mole K	Gas constant (universal)
R_d	≈ 287.04 J/kg K	Gas constant (dry air)
R_v	≈ 1461.50 J/kg K	Gas constant (water vapor)
T_{ref}	≈ 301 K	Constant reference temperature
L_v	$\approx 2.435 \cdot 10^6$ J/kg	Latent heat of vaporization
c_p	≈ 1005 J/K kg	Heat capacity of dry air

Table 1: Constants and parameters for multicloud parameterizations.

Description	Expression
Midlevel θ_{em}	$\theta_{em} = q + \frac{2\sqrt{2}}{\pi}(\theta_1 + \alpha_2\theta_2)$
Precipitation	$P = H_d + \xi_s H_s + \xi_c H_c$
Downdrafts	$D = m_0(1 + \mu(H_s - H_c)/Q_{R01})^+(\theta_{eb} - \theta_{em})$
Radiation	$\text{Rad}_1 = Q_{R01} - \frac{\theta_1}{\tau_D}, \quad \text{and} \quad \text{Rad}_2 = Q_{R02} - \frac{\theta_2}{\tau_D}$
CAPE	$CAPE = \overline{CAPE} + R(\theta_{eb} - \gamma(\theta_1 + \gamma_2\theta_2))$
Lower level CAPE	$CAPE_l = \overline{CAPE} + R(\theta_{eb} - \gamma(\theta_1 + \gamma'_2\theta_2))$
Deep heating	$H_d = \left[\sigma_d \bar{Q} + \frac{\sigma_d}{\bar{\sigma}_d \tau_c^0} (a_1 \theta_{eb} + a_2 q - a_0(\theta_1 + \gamma_2 \theta_2)) \right]^+$
Congestus heating	$H_c = \sigma_c \frac{\alpha_c \bar{\alpha}}{H_m} \sqrt{CAPE_l^+}$

Table 2: Summary of important diagnostic quantities in multicloud model. For more details, the reader is referred to FMK13.

To mimic the behavior of convective variability within a typical GCM grid box, a rectangular $n \times n$ lattice is considered. Each element of the lattice is occupied by a congestus, deep or a stratiform cloud or is a clear sky site. It is represented by an order parameter that takes accordingly the values 0,1,2 or 3. A continuous time stochastic process is then defined by allowing the transitions, for individual cloud sites, from one state to another according to intuitive probability transition rates, which depend on the resolved (“large scale”) variables. These large scale variables are the convective available potential energy integrated over the whole troposphere (CAPE), low level CAPE (see Table 2) and the dryness of the mid troposphere, which are themselves functions of the multicloud model variables such as $\theta_{eb}, \theta_1, \theta_2, q$, etc. For example, the dryness of the mid troposphere is a function of the difference between the atmospheric boundary layer (ABL) temperature θ_{eb} and the middle tropospheric potential temperature θ_{em} . The inclusion of the dryness of the middle troposphere accounts for mixing of the convective parcels with dry environmental air is conceptually similar to the

Table 3: Transition rates and time scales in the stochastic multicloud model simulations. Here $\Gamma(x) = 1 - \exp(-x)$ for $x > 0$ and zero otherwise, $C_l = CAPE_l/CAPE_0$, $C = CAPE/CAPE_0$ and $D = (\theta_{em} - \theta_{eb})/T_0$. For more details, the reader is referred to FMK13.

Description	Transition Rate	Time scale (h)
Formation of congestus	$R_{01} = \frac{1}{\tau_{01}} \Gamma(C_l) \Gamma(D)$	$\tau_{01}=1$
Decay of congestus	$R_{10} = \frac{1}{\tau_{10}} \Gamma(D)$	$\tau_{10}=1$
Conversion of congestus to deep	$R_{12} = \frac{1}{\tau_{12}} \Gamma(C)(1 - \Gamma(D))$	$\tau_{12}=1$
Formation of deep	$R_{02} = \frac{1}{\tau_{02}} \Gamma(C)(1 - \Gamma(D))$	$\tau_{02}=3$
Conversion of deep to stratiform	$R_{23} = \frac{1}{\tau_{23}}$	$\tau_{23}=3$
Decay of deep	$R_{20} = \frac{1}{\tau_{20}} (1 - \Gamma(C))$	$\tau_{20}=3$
Decay of stratiform	$R_{30} = \frac{1}{\tau_{30}}$	$\tau_{30}=5$

switch A in the deterministic multcloud model(KM06a, KM06b,KM07, KM08a, KM08b).

The probability rates are constrained by a set of intuitive rules which are based on observations of cloud dynamics in the tropics (e.g. Johnson et al. (1999); Mapes (2000), KM06a, and references therein). Following KBM10, a clear site turns into a congestus site with high probability if low level CAPE is positive and the middle troposphere is dry. A congestus or clear sky site turns into a deep convective site with high probability if CAPE is positive and the middle troposphere is moist. A deep convective site turns into a stratiform site with high probability. Finally, all three cloud types decay naturally to clear sky at some fixed rate. All other transitions are assumed to have negligible probability. These rules are formalized in Table 3. Notice that the assumption that the transition rates depend on the large scale variables accounts for the feedback of the large scales on the stochastic model, while ignoring the interactions between the lattice sites all together implies that the stochastic processes associated with the different sites are identical (independent and identically distributed). The latter simplification makes it easy to derive the stochastic dynamics for the GCM grid box cloud coverage alone, which can be evolved without the detailed knowledge of the micro-state configuration, by using a coarse-graining technique (Katsoulakis et al., 2003,b) that yields in this case a system of three dimensional birth-death stochastic process for the congestus, deep and stratiform cloud fractions σ_c, σ_d and σ_s respectively (KBM10). The cloud fractions impact the large scale dynamics through algebraic congestus and deep heating closures, listed in Table 2, and dynamical closure for stratiform heating below,

$$\partial_t H_s = \frac{1}{\tau_s} (\alpha_s \sigma_s H_d / \bar{\sigma}_d - H_s). \quad (8)$$

More detailed description of the stochastic multcloud model can be found in FMK13.

2.3. Cloud Radiation Feedback

Clouds play a key role in the understanding of weather and climate. The stochastic multcloud model outlined in the previous section does a remarkable
200 job of capturing the distributions of the three cloud types (congestus, deep and stratiform) that characterize tropical convection in a paradigm two baroclinic mode system. Here we augment the stochastic multcloud model by addition of radiative convective feedback (which will also sometimes be referred to interchangeably as cloud radiation feedback). The impact of radiative convective
205 feedback of each of the three cloud type is parameterized through a product of the cloud fraction and two parameters, which represent idealized projection of the effect onto the two baroclinic modes of the system. The physical motivation for the CRF parameterization is outlined below.

Generally, low and thick clouds, such as congestus, primarily reflect solar
210 radiation and cool the surface of the Earth. The radiative fluxes within the deep convective clouds primarily cancel but result in slight warming. High, thin clouds, such as stratiform clouds, transmit some of the incoming solar radiation; at the same time, they trap some of the outgoing infrared radiation emitted by the Earth and radiate it back downward, thereby warming the surface of
215 the Earth and atmosphere below. In this simple framework we will assume stratiform clouds (due to their relative thinness) transmit half of the short wave radiation of the congestus or deep clouds. The degree of cooling or heating of cloud depends on several factors, including the cloud's altitude, its size, and the make-up of the particles that form the cloud, such as ice and aerosols. Here we
220 are primarily concerned with effect of three idealized paradigm cloud types and their impact upon the two baroclinic modes of the system. We are guided by our intuitive understanding of clouds and observational evidence below.

As noted by Kiehl (1993), observations based on Earth Radiation Budget Experiment (ERBE) satellite data indicate that there is a near cancellation of
225 fluxes between tropical long wave and shortwave cloud forcing in regions of deep convective activity. We define cloud radiative forcing as the difference between the radiative heatings for cloud and clear skies. Positive cloud forcing implies

the presence of the cloud warms the atmosphere relative to a clear sky heating profile. In order to infer the effects of cloud radiative forcing, we consider some
230 observational analyses (Cox and Griffith, 1979a,b).

As shown in the observational analysis of Cox and Griffith (1979a) and in CRM studies (Peters and Bretherton, 2005), within the deep cloud, the long-wave and shortwave heating are both essentially zero. The long-wave cooling at cloud top apparently overwhelms the shortwave warming there, and the result is
235 slight cooling in the first baroclinic mode and more cooling in the upper troposphere than the lower troposphere. For stratiform and deep convective clouds, longwave cloud top cooling is roughly balanced by shortwave cloud top warming during daytime. As a result deep convective clouds warm the lower troposphere relative to the upper troposphere due to longwave cloud base warming in the
240 lower troposphere (Cox and Griffith, 1979a,b). This results in the positive projection onto both first and second baroclinic modes to create a heating profile biased towards lower troposphere. By the same reasoning (Cox and Griffith, 1979a,b; Peters and Bretherton, 2005), stratiform clouds, which follow the deep convection, warm the upper troposphere relative to the lower troposphere due
245 to longwave cloud base warming in the upper troposphere. This corresponds to the choice of heating profile slightly biased towards upper level heating through negative second baroclinic mode contribution. The evidence for effects of congestus cloud forcing is less clear (Cox and Griffith, 1979b). The congestus clouds do not extend in the upper troposphere, and therefore the forcing here is zero.
250 The observational evidence (Schumacher et al., 2004) points to wards cooling of the lower troposphere relative to the upper troposphere by congestus clouds due to longwave cloud top cooling. This corresponds to cooling in of both first and second baroclinic modes by congestus radiative feedback.

The cloud radiative forcing inferred from the observations above is projected
255 on two the baroclinic modes of variability in the model. Formally, we propose the following parameterization for radiative flux terms $S_i, i = 1, 2$ in equations 1 :

$$S_i = -\theta_i/\tau_D + \sigma_{cs}R_i^{clear} + (\sigma_c R_i^c + \sigma_d R_i^d + \sigma_s R_i^s)\alpha, i = 1, 2 \quad (9)$$

Here, $\sigma_{cs} = 1 - \sigma_c - \sigma_d - \sigma_s$, is clear sky fraction. We also introduce tuning parameter, α . Setting $\sigma_{cs} = 1$ and $\alpha = 0$ reduces radiative flux to a simpler formulation, $S_i = \theta_i/\tau_D + R_i^{clear}$, which is use all previous iterations of the multicloud model. As was the case with FMK13, the value of $R_1^{clear} = -1$ K/day is used for first baroclinic mode heating. The value of R_2^{clear} is set by RCE and is on order of half of Kelvin per day. We choose to combine the short wave and longwave cloud radiative effect. In particular, we suppress the role of the short-wave fluxes as their calculation would necessitate complex compositions involving incoming and outgoing radiation and solar zenith angle. Plots of the cloud forcing and cloud radiation heating rate profiles for particular choice of parameters in Table 4 are shown in Figure 1. Left panel of the figure shows idealized convective heating of the three cloud types. Since there is no congestus precipitation in this parameter regime ($\xi_c = 0$), congestus heating projects purely on second baroclinic mode, warming lower troposphere (and cooling upper troposphere). Deep convection projects only on the first baroclinic mode, warming the mid-troposphere. Stratiform rain projects onto the first baroclinic, through, $\xi_s = 0.5$, in addition to warming the upper troposphere and cooling (through evaporation) of lower troposphere. This results in the profile skewed towards the upper troposphere heating. The right panel shows schematics of cloud radiative forcing profile computed from each cloud type and clear sky. The schematic is computed by imposing values in Table 4 on the two baroclinic modes and (for ease of comparison) assuming that each cloud completely covers the sky. The cloud forcing profiles are similar to the ones used by Schumacher et al. (2004) (see Figure 11 of that study) to process Tropical Rainfall Measuring Mission (TRMM). The formulation above is also based on the GATE (GARP Atlantic Tropical Experiment) data studies (Cox and Griffith, 1979a,b). The clear sky cooling rate is comparable to the values used in previous multicloud model studies. Lastly, this type of direct CRF parameterization, where CRF is assumed to be proportional to cloud fraction or heating (often through precipitation efficiency parameter), is commonly a part of the radiation scheme in both simple one baroclinic mode model (Sobel et al., 2004) and CRM studies

(Bony and Emanuel, 2005; Zurovac–Jevtik et al., 2005).

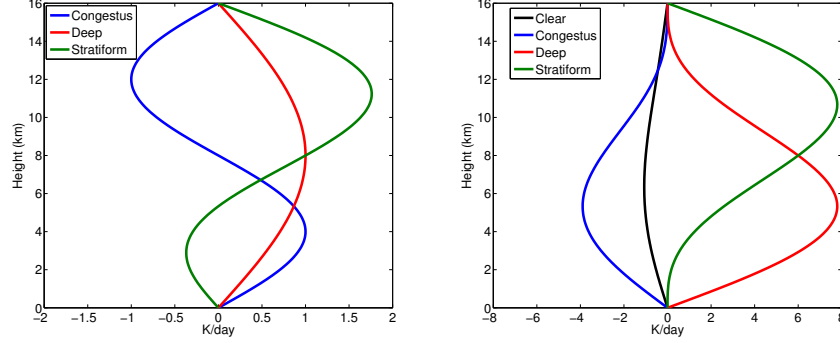


Figure 1: Left: Idealized convective heating of the three cloud types. Since there is no congestus precipitation in this parameter regime ($\xi_c = 0$), congestus heating projects purely on second baroclinic mode. Deep convection projects only on the first baroclinic mode. Stratiform rain projects onto the first baroclinic (through $\xi_s = 0.5$) in addition to warming the upper troposphere, thus resulting in profile skewed towards the upper troposphere heating. Right: schematics of cloud radiative forcing profile computed from each cloud type and clear sky. The schematic is computed by imposing values in Table 4 on the two baroclinic modes and (for ease of comparison) assuming that each cloud completely covers the sky. The clear sky cooling rate is of profile typically used in the multcloud models.

290 2.4. Slab mixed-layer ocean

We use a simple model for an interactive bulk mixed-layer ocean. The character of our mixed-layer ocean will be subsumed into a single variable, T_s , the sea surface mixed layer temperature. The evolution of T_s will be given by

$$c_0 \frac{dT_s}{dt} = S - \frac{c_0}{\tau_{eo}} (\theta_{eb}^*(T_s) - \theta_{eb}) \quad (10)$$

where c_o is the heat capacity for the mixed layer ocean. The radiation flux, $S = S^{rad} + S^{out}$, includes radiative effects, S^{rad} and an imposed ocean heat transport, S^{out} . The value of the ocean layer evaporation constant, τ_{eo} is determined by RCE conditions, and the surface temperature variable T_s is an anomaly from equilibrium conditions (i.e., it takes the value of zero in RCE).

Table 4: Parameters and constants used for the cloud radiative feedback and ocean coupling. Note that $R_2^{clear} \approx 0.3$ K/Day is determined by RCE and is comparable to the value used in the previous multcloud model studies.

Parameter	CRF rate coefficient associated with	value
R_1^c	Congestus clouds for 1st baroclinic mode	-3 K/day
R_1^d	Deep clouds for 1st baroclinic mode	6 K/day
R_1^s	Stratiform clouds for 1st baroclinic mode	6 K/day
R_1^{clear}	Clear sky for 1st baroclinic mode	-1 K/day
R_2^c	Congestus clouds for 2nd baroclinic mode	-1.5 K/day
R_2^d	Deep clouds for 2nd baroclinic mode	3 K/day
R_2^s	Stratiform clouds for 2nd baroclinic mode	-3 K/day
R_2^{clear}	Clear sky for 2nd baroclinic mode	-0.2 K/day

The sensible heat flux will be ignored, since its contribution to the energy bud-
 300 get is small compared to the shortwave radiation and latent heat fluxes (Peters
 and Bretherton, 2005; Sobel et al., 2004; Kikuchi and Wang, 2008; Frenkel et al.,
 2010).

In their model for Walker circulation, Sobel et al. (2004) used a similar
 mixed-layer ocean model. The authors used zonally varying term for the com-
 305 bined effect of ocean heat transport and clear sky shortwave forcing. Also, a
 similar model with zonally varying ocean heat transport has been used in Peters
 and Bretherton (2005). Here, we omit explicit representation of the ocean heat
 transport and instead use equation 7 to account for zonal SST variation im-
 plicitly, through E . This facilitates a direct comparison to previous multcloud
 310 studies.

As shown in Peters and Bretherton (2005); Sobel et al. (2004), the long-wave
 cooling of the surface is nearly cancelled by back radiation from the atmosphere.
 Due to the smallness of longwave cloud forcing we will not consider long-wave
 radiation for cloud forcing at the surface, but will include long-wave effects in
 315 the constant clear sky flux. For the radiative flux S^{rad} at the surface, we will
 have a clear sky component S^{clear} and a cloud forcing component S^{cf} :

$$S^{rad} = S^{clear} + S^{cf} \quad (11)$$

For clear sky, the net shortwave flux averaged over one day is roughly
 $S^{clear} = 300 \text{ W/m}^2$, as shown in Sobel et al. (2004); and the net long-wave
 flux is roughly $S_{lw}^{clear} = -50 \text{ W/m}^2$, as shown in Peters and Bretherton (2005).
 320 Therefore, we will choose the daily-averaged clear sky flux to be

$$S^{clear} = S_{lw}^{clear} + S_{sw}^{clear} \quad (12)$$

For the radiative cloud forcing, we use

$$S^{cf} = -S_{sw}^{clear}(\sigma_c + \sigma_d + 0.5\sigma_s) \quad (13)$$

In the equation above, we make an assumption that stratiform clouds, due to
 their relative thinness, block half as much short wave radiation compared to the

thicker congestus and deep clouds. However, a large area fraction of stratiform
 325 clouds makes them extremely important for radiative feedback effects.

We can further rewrite equation 10 in an explicit form.

$$\frac{dT_s}{dt} = \frac{S^{clear}(1 - A^c) - S^{sw}A^c}{c_0} - \frac{1}{\tau_{eo}}(\theta_{eb}^*(T_s) - \theta_{eb}) \quad (14)$$

Here we use the sum of cloud area fractions to determine fraction cloud cov-
 erage of the sky $A^c = (\sigma_c + \sigma_d + 0.5\sigma_s)\alpha_o$, along with a tuning parameter α_o .
 This tuning parameter is similar to the tuning parameter α used for radiative
 330 feedback coupling in equation 9. The corresponding radiative factor for convec-
 tive clouds is given by $S^{sw} \approx 200 \text{ W/m}^2$, as estimated above. The mixed layer
 heat capacity c_0 is given by

$$c_0 = c_{R,0}\rho_0 h_{ml} \quad (15)$$

This calculation is based on (heat capacity) $c_{R,0} = 4000 \text{ J/kg K}$ and density ρ .
 When depth of mixed layer, h_{ml} , is set to 20 meters, $c_0 \approx 8 \times 10^7 \text{ JK}^{-1}\text{m}^{-2}$. This
 335 mixed layer depth corresponds to approximately 40 day time scale for mixed
 ocean layer. In select simulations presented here, we will vary mix layer depth,
 by setting $h_{ml} = 10, 20$ and 40 meters, which results in AOC time scale of
 20, 40 and 80 days, respectively, for mixed layer. The ocean layer T_s is coupled
 to the atmospheric boundary layer (θ_{eb}) through atmospheric boundary layer
 340 saturation equivalent potential temperature. A simple closure, $\theta_{eb}^*(T_s) = 5T_s$,
 is derived from the Clausius–Clapeyron relation as shown in Appendix A. In
 simulations without ocean, saturation equivalent potential temperature, θ_{eb}^* , is
 independent of T_s .

3. Single column simulations

345 In this section, the effects of the new mechanisms of cloud-radiation feedback
 and atmosphere-ocean coupling, are studied in the context of single column
 simulations and compared to the FMK13 results. The single column equations
 are obtained by disregarding spatial dependence components and the zonal wind.

As in KBM10, we employ a third order Adams-Bashforth method to integrate
 350 the dynamical core ODEs. The coarse grained birth-death process is evolved
 in time by means of Gillespie’s exact algorithm (Gillespie, 1975, 1977). All
 simulations in this section are run for 2000 days, while a 10 or 100 days interval
 of the solution is shown.

To facilitate the comparison, we first review the basic results of FMK13.
 355 Figure 2 show the stochastic simulation of FMK13. The most notable feature is
 the time synchronization of the oscillations of the stochastic and deterministic
 variables which leads to time series with frequent precipitation peaks of 10
 K/Day and more intermittent large precipitation events on the order of 20
 K/Day. Each convective event is initialized by a build up of low level CAPE.
 360 The resulting congestus clouds moisten the atmosphere. This moist atmosphere,
 combined with the build up of CAPE, produces deep convective events which are
 in turn followed by stratiform clouds. The relationship between small and large
 precipitation events is reminiscent of a progressive deepening of convection on
 multiple scales (Mapes et al. 2006). By design, the congestus clouds are followed
 365 by deep convective and trailing stratiform clouds. The transitions rates are
 associated with moisture and dryness. In particular, we see a high correlation
 between positive, θ_{el} and moisture anomalies and deep convective activities. In
 this simulation, the ocean mixed layer temperature equation is slaved to the
 atmospheric variables and does not feedback into the model dynamics (which
 370 are driven under the assumption that T_s is actually fixed at its equilibrium
 value).

Figures 3 shows the time series of simulations with cloud radiation feedback
 and atmosphere ocean layer and with radiative convective feedback. We choose
 small value of CRF strength ($\alpha = 0.1$) and 40 meter deep ocean layer. The
 375 top panel shows the contribution of the radiative feedback to the heating. The
 effects of radiative convective feedback are subtle. The feedback accounts for
 roughly one tenths of the heating in the column, which most significant contri-
 bution coming from large deep convective events and trailing stratiform anvils.
 On the other hand, the atmosphere ocean coupling produces pronounced en-

380
velope of convective activity with roughly 30 day period of oscillation (as will be shown in the following figures), associated with the ocean temperature fluctuations (bottom panel). The clear sky conditions lead to increase in mixed ocean layer temperature, which leads to increase in ABL temperature anomalies, and creation of deep convection and stratiform anvils, which cools mixed
385
ocean layer, in turn, leading to smaller convective fraction and clear sky, closing the loop. Overall, the dynamics of the model becomes more irregular with introduction of ocean coupling and CRF. In fact the model has suppressed and enhanced convection periods associated with time scale related to the ocean depth. The direct link between the ocean layer depth and the intraseasonal
390
and subannual oscillations will be discussed below in more detail. Briefly, the oceanic dynamics introduces a new time scale, τ_{eo} , into the system, and this atmosphere-ocean coupling time scale is relatively slow and introduces low frequency oscillations. These oscillations resemble suppressed and active phases of MJO but on different time scales. We note that, just like in MJO, the sup-
395
pressed phase of oscillation has intermittent deep convection and small average cloud fraction, while the active phase of oscillation consists of highly organized coherent convective episodes.

400
In Figure 4 we study the response of the model to variation in the strength of the CRF in presence of the 20 meter deep ocean layer coupling. For reference, panel A shows results from FMK13 (see Figure 3 in FMK13). The mixed ocean layer is added in simulations of panel B, we note that ocean coupling introduces weak low frequency envelope to the time series. It can be seen that the period of oscillation is on order of 30 days. The oscillation does not seem to be associated with a deterministic instability since, in a linear stability analysis (not shown) of a deterministic version of the model, there is a stable RCE state. The addition of weak CRF coupling ($\alpha = 0.1$), in panel C, enhances this effect but shortens the frequency of the oscillation. It also makes oscillations in the ocean mixed layer more chaotic. The strong CRF coupling ($\alpha = 0.4$), of panel D, destroys the low frequency envelope. The values of α higher than approximately 0.4 produce unphysical large climatology. This is somewhat similar to, but not as extreme
410

as, the findings of Sobel et al. (2004), where strong values of the convective feedback tuning parameter were seen to cause solutions to blow up.

In Figure 3, we experiment with the depth of the ocean mixed layer, while keeping the cloud radiation feedback constant (at weak value of $\alpha = 0.1$). It
415 appears that the mixed ocean layer envelope modulation is prominent for all ocean depths considered. To quantify the changes in the behavior we study the Fourier spectrum of the precipitation in Figure 5. Firstly, we observe that ocean coupling introduces low frequency oscillation. It could further be established that the time scale of this oscillation is proportional to the depth of the
420 layer. Secondly, we draw the conclusion that strong convective radiative feedback interferes with the low frequency introduced by the ocean mixed layer. The stronger values of radiative feedback lead to shift of the power in the spectrum from low frequency oscillations associated with ocean layer to higher frequency intrinsic variability. While the results are omitted for conciseness, in absence of
425 ocean coupling, radiative feedback slightly shifts the power spectrum towards the modes with frequency slightly lower than the intrinsic variability of the system but shorter than frequencies associated with bulk ocean layer. The Fourier spectrum analysis confirms observations made from Figure 4.

It is believed that both CRF and AOC play an important role in low fre-
430 quency tropical variability, such as MJO and ENSO (Bony and Emanuel, 2005). However, it is hard to infer the exact effects of CRF and AOC on these multi-scale phenomena from column simulations. In order to investigate the behavior of the model further we proceed to the spatially extended simulations.

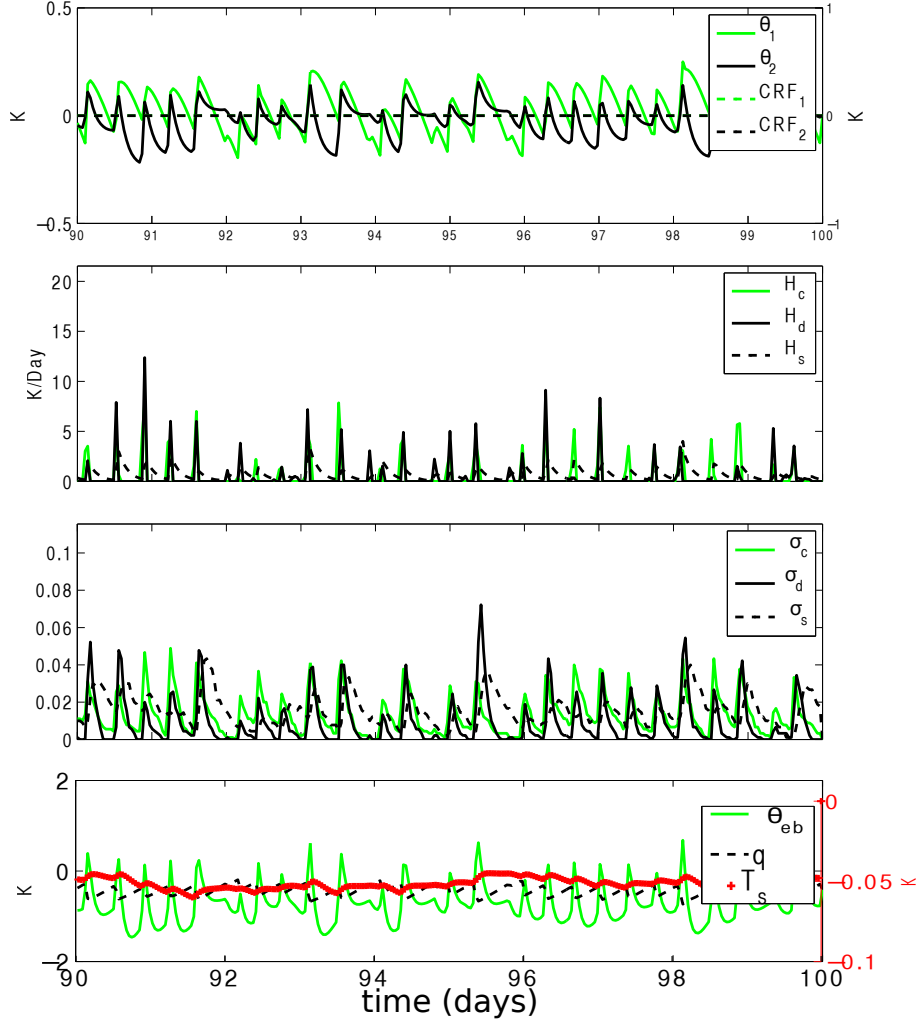


Figure 2: Time series of dynamic variables for FMK13. In this simulation, the ocean mixed layer temperature equation is a passive slave to the system and does not feedback into the model dynamics, and cloud-radiative feedback was not included and hence is plotted with amplitude 0. CRF1 and CRF2 refer to the cloud-radiative feedback portion of the radiative heating S_1 and S_2 , respectively, of (9).

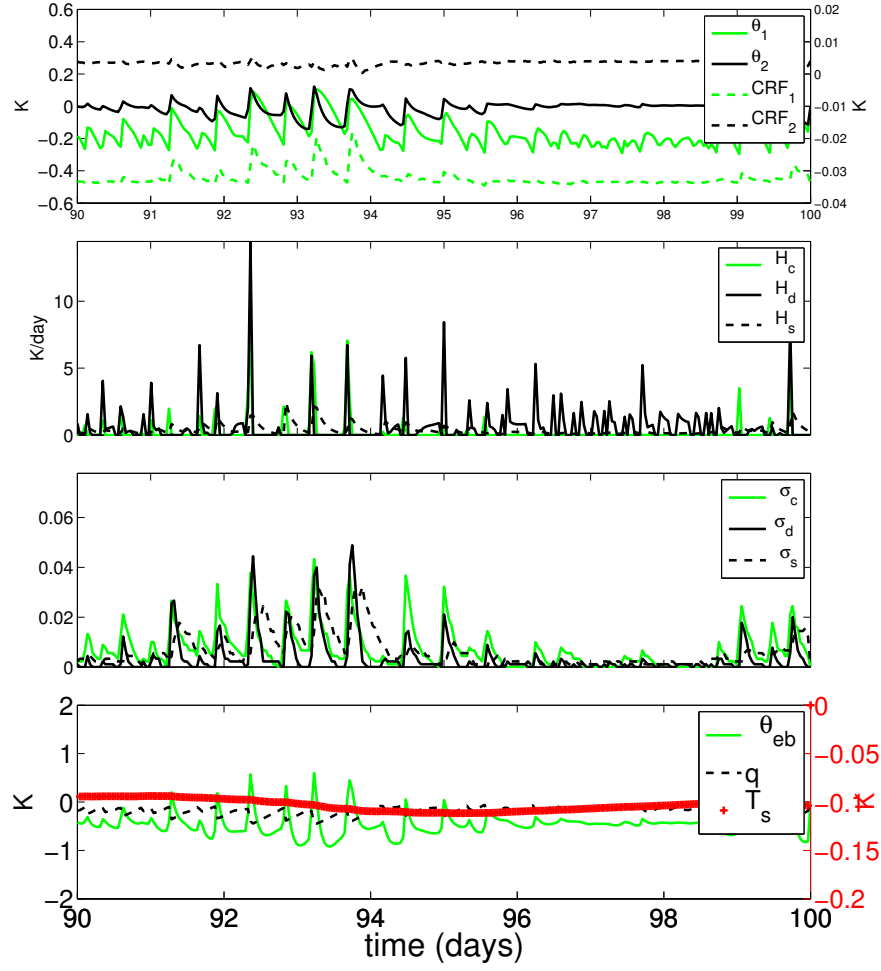


Figure 3: Same as Fig. 2, except with additions of cloud-radiation feedback with $\alpha = 0.1$ and 40 meter deep ocean layer. The effects of cloud-radiation feedback (CRF_1 and CRF_2) are roughly one-tenth the magnitude of the convective heating (H_c, H_d, H_s). Overall, the dynamics of the model is greatly improved by the introduction of the suppressed and enhanced convection periods associated with low frequency sea surface temperature oscillations.

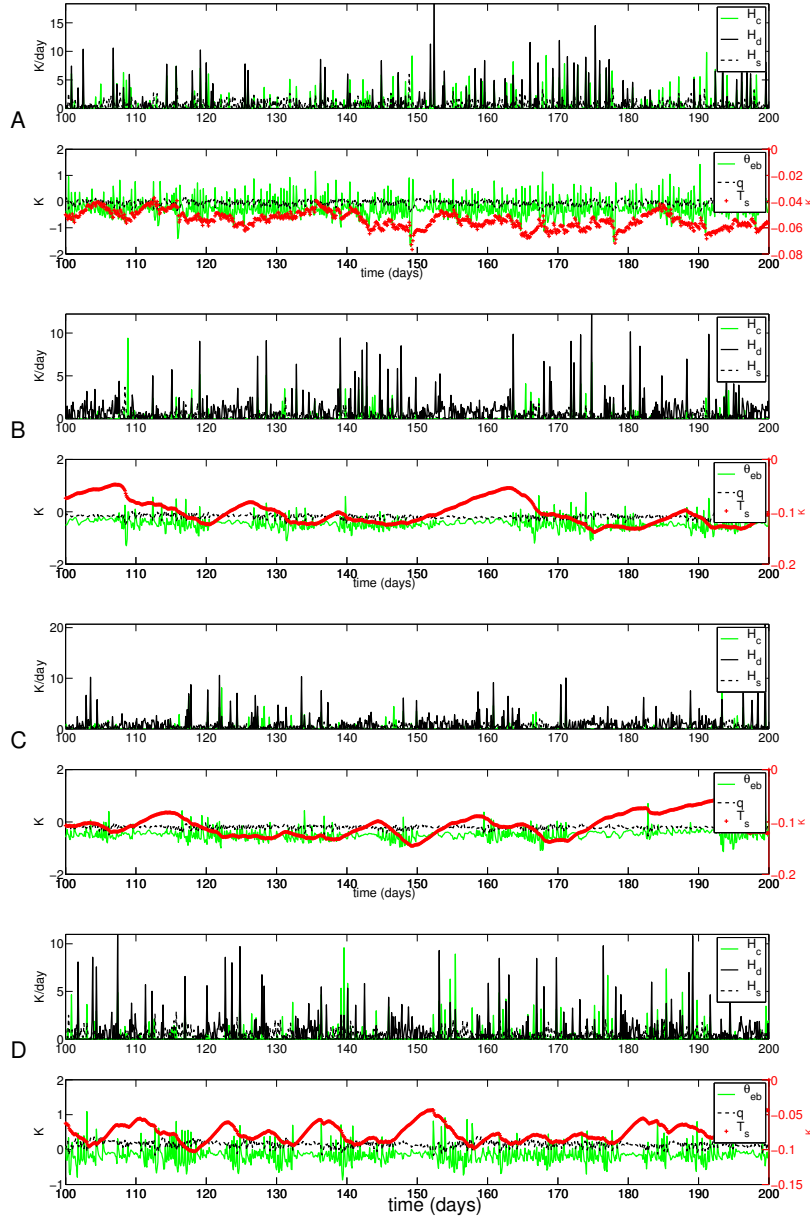


Figure 4: Precipitation and mixed ocean layer temperature for simulations (top to bottom) associated with: A) FMK13, B) simulations with atmosphere ocean coupling (20 meter) and without atmospheric cloud-radiative forcing $\alpha = 0$, C) 20 m ocean with weak feedback $\alpha = 0.1$ and D) 20 m ocean with strong radiative feedback $\alpha = 0.4$.

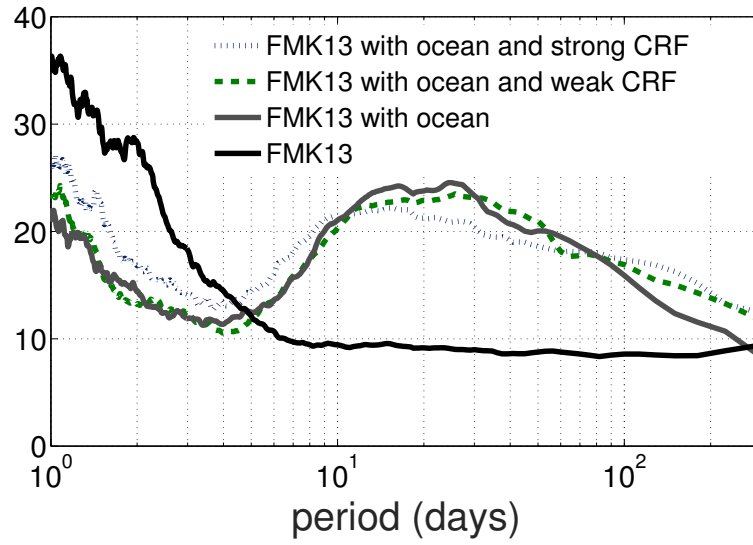


Figure 5: Power spectrum of precipitation time-series for column simulations with FMK13 (solid black), FMK13 with atmosphere-ocean coupling (20 meter) and without radiative feedback $\alpha = 0$ (solid grey), FMK13 with atmosphere-ocean coupling (20 meter) and with weak radiative feedback $\alpha = 0.1$ (green dash) and FMK13 with atmosphere-ocean coupling (20 meter) and with weak radiative feedback $\alpha = 0.4$ (fine blue dash). This corresponds to four cases shown in Figure 4.

4. Spatially extended simulations with a warm pool

435 This section presents the results of spatially extended simulations for the stochastic multcloud model with non-uniform SST backgrounds, mimicking the Indian Ocean western Pacific warm pool (Section 2.4). Simulations with new mechanisms are considered and compared to FMK13, CRMs, and observations. We first consider separately variation of CRF and ocean layer depth before
440 combining the effects. The results are summarized in Table 5.

The numerical method used is an operator-splitting strategy where the conservative terms are discretized and solved by a non-oscillatory central scheme while the remaining convective forcing terms are handled by a second-order Runge-Kutta method (Khouider and Majda, 2005a,b). As for the single column
445 simulations, the stochastic component of the scheme is resolved using Gillespie’s exact algorithm (Gillespie, 1975). We consider the same parameter regimes discussed in the previous section and perform 800 day simulations, with a 5 minute time step and a resolution of 40 km.

4.1. Variation of either CRF or ocean layer depth

450 In all we consider six distinct model configurations. In addition to FMK13, we present simulations with weak and strong CRF ($\alpha = 0.1$ and 0.4 respectively). Separately, we consider addition of 10,20 and 40 meter deep ocean layer to FMK13. Figure 6 shows mean zonal vertical structure for all the regimes computed from 2000 days of data. We note that in all cases, an introduc-
455 tion of ocean layer decreases the strength of the zonal mean winds and makes the relative strength of the first baroclinic mode component stronger. The radiative convective feedback allows for slightly stronger circulation with higher second baroclinic component. For warm pool simulations, introduction of CRF increases strength and mean of the Walker cell by about 10 percent. The re-
460 sults are summarized in Table 5. The significant result here is low level cooling associated with the ocean coupling. The ocean acts as an energy sink and contributes to the weaker mean Walker circulation. This is in sharp contrast

CRF	Ocean	SST	(U ,W)	std(u_1)
α			(m/s,cm/s,)	m/s
0	-	5K	(9.3 ,0.86)	0.18
0.1	-	5K	(9.7 ,0.97)	0.19
0.4	-	5K	(10.5 ,1.11)	0.19
0	10 m	5K	(8.0 ,0.63)	0.18
0	20 m	5K	(7.7,0.62)	0.17
0	40 m	5K	(7.6 ,0.61)	0.14
0.4	40 m	5K	(8.9,0.7)	0.20
0	-	2.5K	(4.0 ,0.53)	0.13
0	40 m	2.5K	(3.9 ,0.50)	0.23
0.4	40 m	2.5K	(4.3,0.46)	0.20

Table 5: Mean and variability of the Walker circulation in the spatially extended simulations. Each case uses different values of the cloud–radiation feedback parameter α , ocean layer depth, and amplitude of the warm pool sea surface temperature. For the time-averaged Walker circulation, the statistics reported are the maxima of U and W and the standard deviation of u_1 .

with counterintuitive result of Sobel et al. (2004), where the authors propose recharge-discharge theory to account for an increased heating associated with AOC. On the other hand anomaly effects on wind evaporation feedback are
465 completely absent in the present models. Also, since only the mixed ocean layer instead of ocean dynamics is included in this work, the cloud over the warm pool will reduce the shortwave radiation and cool the SST.

The Figure 7 shows deviations from zonal mean velocity field for all the
470 regimes. We note that the simulations with the CRF produce the strongest intermittent bursts of convection. In particular the both weak and strong CRF ($\alpha = 0.1$ and 0.4 respective) simulations in panels B and C are more intermittent than the FMK13 simulation in panel A, in the sense that panel A displays a repetitive sequence of convectively coupled wave events, each of roughly the
475 same amplitude and repeating roughly every 15 days, whereas panels B and C display irregular periods of time between convectively coupled wave events and a mixture of periods of weak convectively coupled wave activity (such as times 870–890 days in panel B and times 700–740 days in panel C) and strong convectively coupled wave activity (such as times 830–850 days in panel B and
480 times 860–890 days in panel C). These simulations also produce the strongest variability and mean, as shown in Table 5. The simulation with ocean (but without CRF), shown in figure D, shows an interesting low frequency variability on order of 100 days (i.e., periods of strong convectively coupled wave events at times 760–800, 870–890, and 920–940 days are separated by periods of relatively
485 weak convectively coupled wave activity at times 700–760, 800–870, and 890–920 days). The remnants of this behavior can be seen when a weak CRF is also introduced in panel E, but disappear with the stronger CRF of panel F.

Figure 8 shows Fourier spectrum for velocity, atmospheric moisture, ABL and sea surface temperature. We note that variability for the base FMK13
490 case, is characterized by 15 day convectively coupled waves (while the moisture q tends to have maximum power at a time scale of 7 or 8 days). These are associated with synoptic systems originating inside of the warm pool. A similar oscillation is observed by Slawinska et al. (2014b) in a CRM study. The shallower

ocean layer shifts the variability towards (particularly in atmospheric and ABL
495 moisture) lower frequency time scales. A similar effect is associated with lower
SST gradient discussed in the following section. In all cases, the ocean layer
decreases the strength of the variability and introduces low frequency envelope
structures in atmospheric and ocean boundary layer fields. The Fourier analysis
in Figure 8 shows a progressive shift towards the lower frequencies associated
500 with the increase in ocean depth. The same figures shows the effects of the CRF,
which are mixed in nature. The radiative convective feedback allows for slightly
larger amplitude variability but negatively impacts low frequency oscillations in
the system.

Note that, we omit spatially extended homogeneous SST simulations, since
505 it is hard to distinguish (visually) the difference between standard FMK13 sim-
ulations and the ones with low and moderate radiative convective feedback. We
also performed simulations with wind induced surface heat exchange (WISHE)
(Majda and Shefter, 2001). Likewise, these results are also omitted, since the
addition of WISHE mechanism did not appear to a significant impact on the
510 mean or the variability of the stochastic multcloud model simulations.

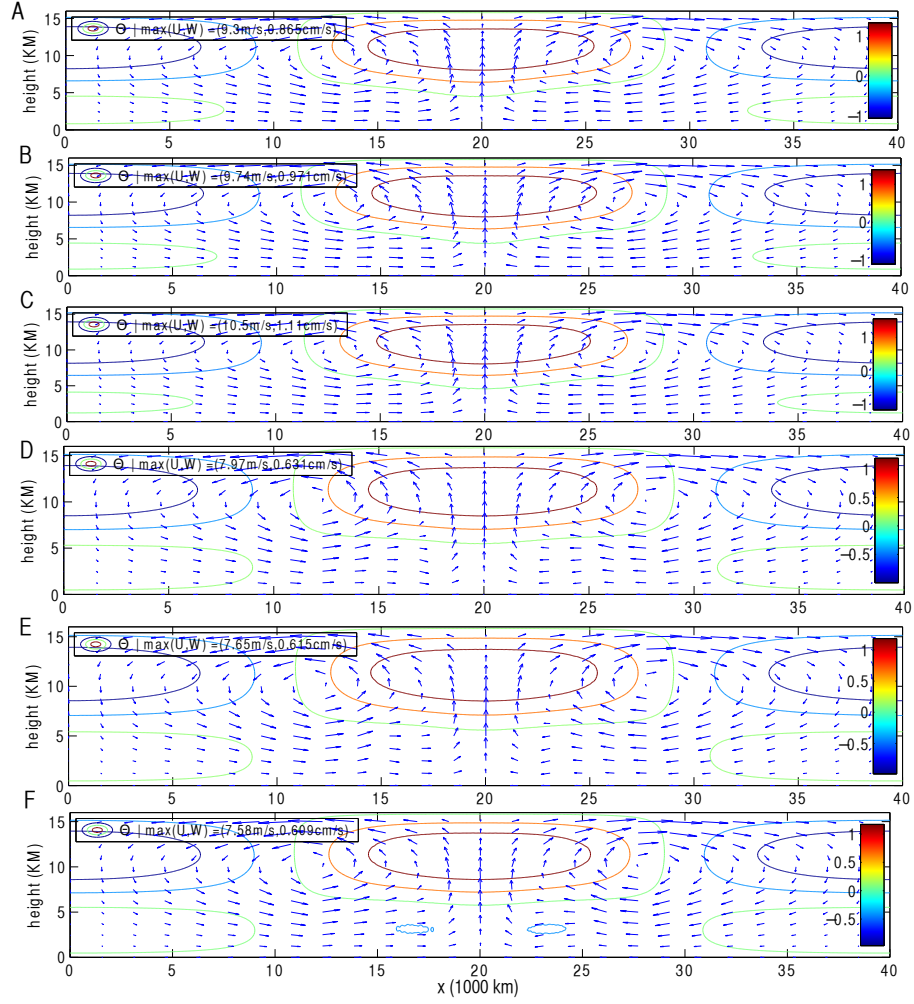


Figure 6: Mean zonal vertical structure associated with: A) FMK13 B) FMK13 with weak CRF ($\alpha = 0.1$) C) FMK13 with strong CRF ($\alpha = 0.4$), D) FMK13 with 10 meter ocean , E) FMK13 with 20 meter ocean, F) FMK13 with 40 meter ocean .

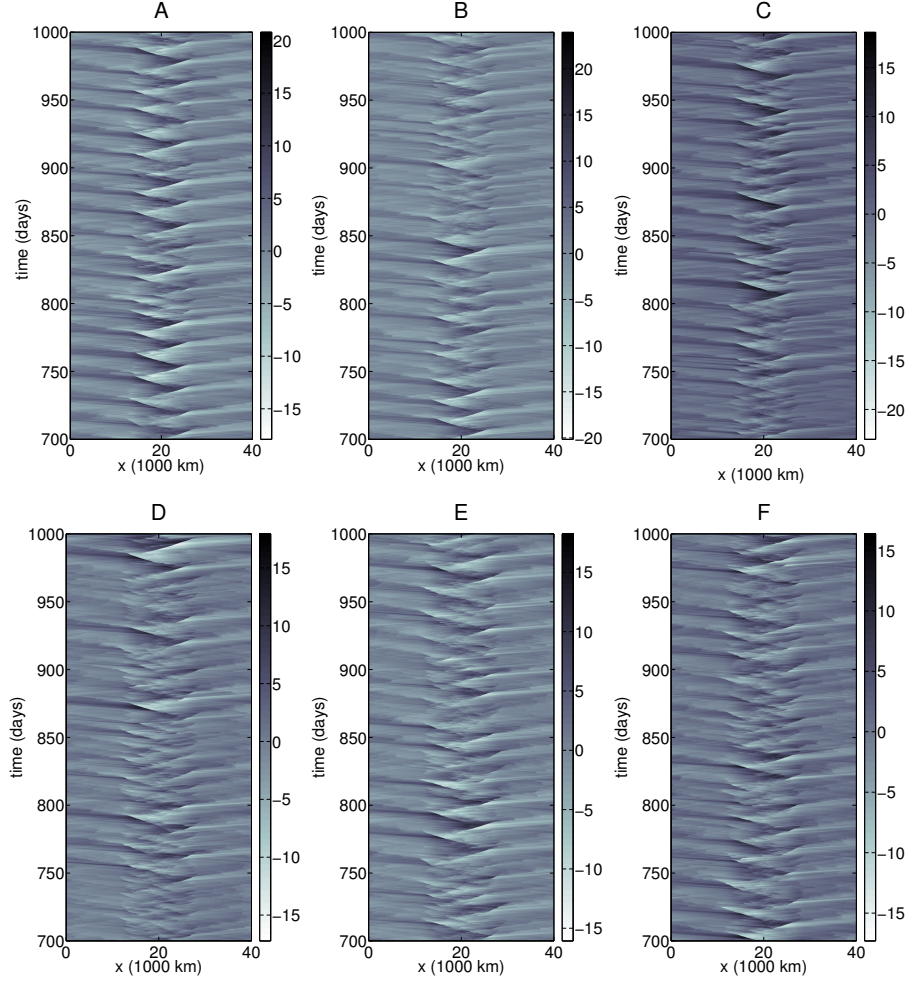


Figure 7: Contours of velocity anomalies $u_1(x, t)$ for: a) FMK13, b) FMK13 with weak CRF ($\alpha = 0.1$), c) FMK13 with strong CRF ($\alpha = 0.4$), d) FMK13 with 10 meter ocean, e) FMK13 with 20 meter, f) FMK13 with 40 meter ocean. Anomalies are computed as deviations from the time-averaged mean state. Note that slightly different color bars are used in different panels.

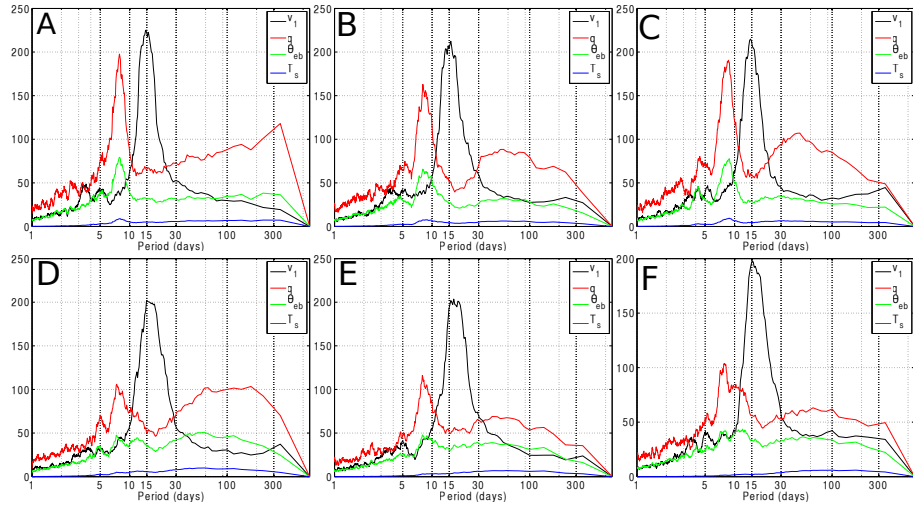


Figure 8: Power spectrum of velocity (black), moisture (red), θ_{eb} (green), and sea surface temperature (blue) from 500 days of observations at 20 equally spaced sites in the warm pool for simulations associated with a) FMK13, b) FMK13 with weak CRF ($\alpha = 0.1$), c) FMK13 with strong CRF ($\alpha = 0.4$), d) FMK13 with 10 meter ocean, e) FMK13 with 20 meter ocean, f) FMK13 with 40 meter ocean. Note that a slightly different y -axis range is used in panel (f).

4.2. Combining effects of cloud radiative feedback and atmosphere- ocean coupling

Here we consider three distinct model CRF/AOC configurations :FMK13, FMK13 with ocean coupling , FMK13 with ocean coupling and weak $\alpha = 0.1$ radiative feedback. In addition, we use warm pool strength of 5K (typical of multicloud models) and 2.5 K for each of three model configurations. Thus in total we consider six total parameter regimes, which highlight the combined effects of CRF and AOC, as well as SST strength.

To begin, we consider in the detail the dynamics of the spatially extended model with weak CRF and 40 meter ocean layer coupling in Figure 9. The contours of velocity field (upper left panel) show three large convective events in the center of the warm pool occurring in span of 50 days. The events are associated with build up of boundary layer moisture (lower left panel) which in turn appears to be coupled with the ocean temperature (middle left panel). These large convectively coupled waves have a high deep convective heating coupling and moisture content. They also produce faster convectively coupled gravity waves outside of the warm pool. These waves carry less moisture (lower right panel) but still produce intermittent deep convective events in the suppressed regions of the Walker circulation. We also observe a myriad of small amplitude standing wave activity in θ_{eb} and T_S fields inside of the warm pool. These are associated with the congestus clouds that are responsible for the moistening of the mid-troposphere and preconditioning the system for the next convective event. The congestus heating is generally abundant inside of the positive SST area of the warm pool.

The Figure 10 show mean zonal vertical structure for all the regimes computed from 2000 days of data. We note that in all cases, ocean layer decreases the strength of the mean and make for more first baroclinic mode dominated circulation. The radiative convective feedback allows for slightly stronger circulation with higher second baroclinic component. For warm pool simulations, introduction of CRF increases strength and mean of the Walker cell by about 10 percent. Unsurprisingly, weaker SST gradient results in weaker mean circula-

tion. The effects of ocean coupling are also less drastic in the weak SST gradient case. In fact, the simulation with ocean and CRF mechanism is stronger than the FMK regime for 2.5K SST forcing.

545 More importantly, the weaker SST gradient fundamentally changes the frequency of the generation of the synoptic scale waves inside the warm pool. Figure 11 shows deviations from zonal mean velocity field for all the regimes. The weaker 2.5K warm pool shifts the variability towards 30 day convectively coupled waves which have moisture and ABL temperature spectral peak near
550 150 days. Unlike the mean, the amplitude of the waves is not impacted by weaker SST gradient. The simulations with ocean and CRF mechanisms appear more intermittent compared to FMK13 simulations. A remnant of the low frequency envelope can be seen in the ocean coupled simulations of panel B. This pattern is hard to discern visually when CRF is added in panel C.

555 In order to elucidate the low frequency dynamics in the system, we consider the Fourier spectrum for velocity, atmospheric moisture, ABL and sea surface temperature in Figure 12. For both weak and strong SST gradient, the addition of ocean shifts the power spectrum towards the seasonal time scales. The results are mostly seen in ABL and atmospheric moisture while changes in velocity
560 spectrum (as well as precipitation) are more subtle. Conversely, the introduction of the CRF shifts the spectrum towards marginally higher frequencies (moisture spectrum in particular). The weaker SST gradient fundamentally alters the moisture spectrum: from bimodal to a single low frequency peak. This change in the moisture power spectrum likely indicates slower moisture build up inside
565 of the warm pool and is responsible for the doubling of the period of generation of the synoptic scale disturbances inside of the warm pool.

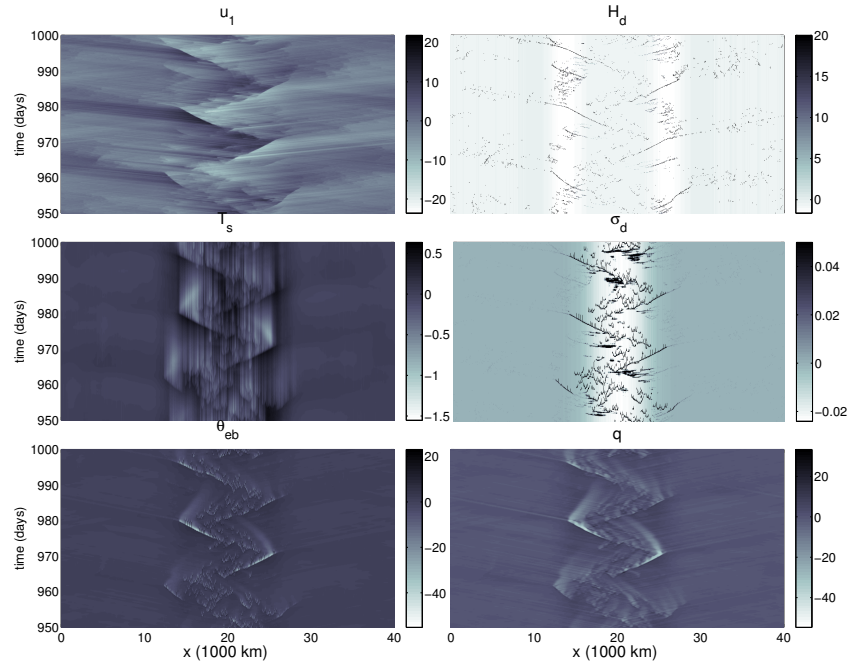


Figure 9: Hovmoller diagram showing 50 days of anomalies for stochastic multicloud model simulation with weak CRF and 40 meter deep ocean layer. The units of the variables are u_1 [m/s], H_d [K/d], T_s [K], σ_d [unitless], θ_{eb} [K], q [K].

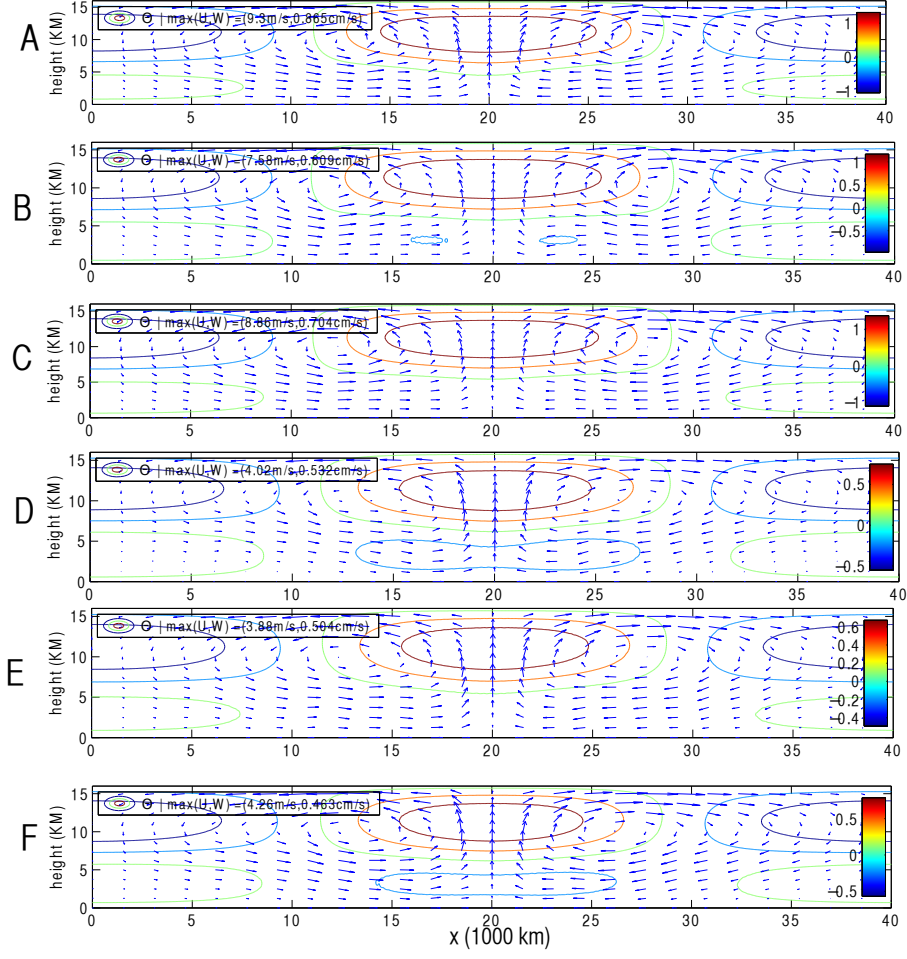


Figure 10: Mean zonal-vertical structure associated with: A) FMK13 (5K warm pool), B) FMK13 with ocean coupling (5K warm pool), C) FMK13 with ocean coupling and radiative convective feedback (5K warm pool), D) FMK13 (2.5K warm pool), E) FMK13 with ocean coupling (2.5K warm pool), F) FMK13 with ocean coupling and radiative convective feedback (2.5K warm pool). Wherever appropriate, we use strong CRF and 20 meter ocean.

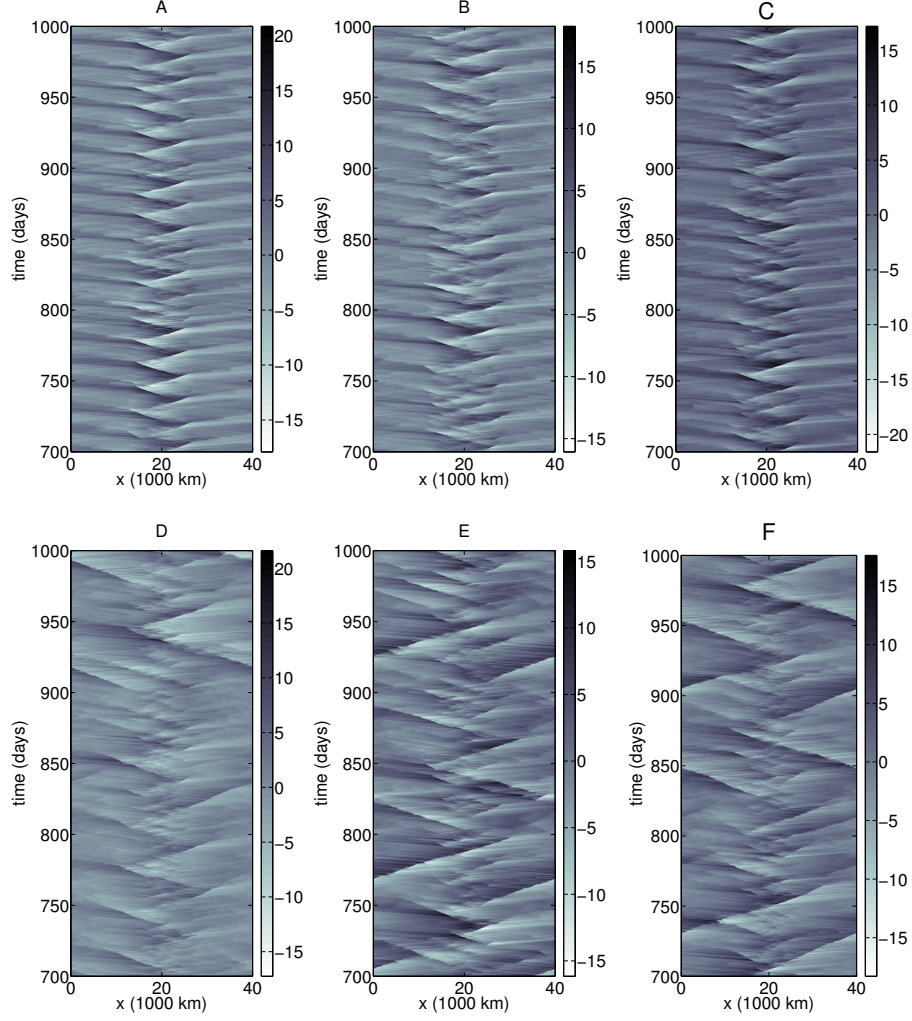


Figure 11: Contours of velocity anomalies $u_1(x, t)$ for an interval of 300 days for: A) FMK13 (5K warm pool), B) FMK13 with ocean coupling (5K warm pool), C) FMK13 with ocean coupling and radiative convective feedback (5K warm pool), D) FMK13 (2.5K warm pool), E) FMK13 with ocean coupling (2.5K warm pool), F) FMK13 with ocean coupling and radiative convective feedback (2.5K warm pool). Wherever appropriate, we use strong CRF and 20 meter ocean. Anomalies are computed as deviations from the time-averaged mean state.

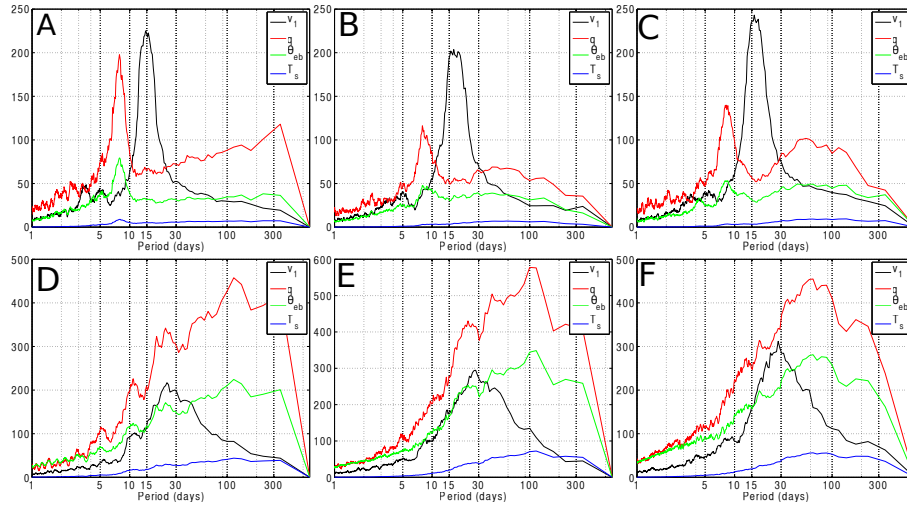


Figure 12: Power spectrum of velocity (black), moisture (red), θ_{eb} (green), and sea surface temperature (blue) from 500 days of observations at 20 equally spaced sites in the warm pool for simulations associated with: A) FMK13 (5K warm pool), B) FMK13 with ocean coupling (5K warm pool), C) FMK13 with ocean coupling and radiative convective feedback and 5K warm pool, D) FMK13 (2.5K warm pool), E) FMK13 with ocean coupling and 2.5K warm pool, F) FMK13 with ocean coupling and radiative convective feedback 2.5K warm pool. Wherever appropriate, we use strong CRF and 20 meter ocean.

5. Conclusions

Here, the stochastic multcloud model modified with new parameterized cloud radiation feedback and atmosphere-ocean coupling mechanisms is used to study horizontally homogeneous one column model dynamics and flows above the equator without rotation effects. The stochastic model is based on a coarse grained Markov chain lattice model where each lattice site takes discrete values from 0 to 3 according to whether the site is clear sky or occupied by a congestus, deep or stratiform cloud. The convective elements of the model interact with each other and with the large scale environmental variables through CAPE and middle troposphere dryness. The multcloud parameterization greatly improves the representation of the cloud distributions compared to the Betts-Miller scheme which typically are used for radiative feedback studies in coarse vertical resolution models (Zurovac-Jevtik et al., 2005; Moncrieff and Klinker, 1997) and allows for results which rival the much more computationally complex GCM simulations (Zurovac-Jevtik et al., 2005). The emphasis in this study is placed on elucidating the role of intraseasonal and seasonal variability in the system.

In Sections 2.3 and 2.4, we propose simple parameterized closures for cloud radiation feedback and atmosphere ocean coupling. The basis for the CRF is an assumption that each cloud has a fixed radiative profile, with strength proportional to the cloud fraction in each computational grid point. Similarly, convective cloud fractions impact the ocean layer temperature by blocking a proportional amount of short-wave solar heating. In single column simulations, the atmosphere-ocean coupling introduces a new slow time scale into the system, which thereby induces low frequency variability, the scale of which varies with ocean mixed layer depth. These intraseasonal and subannual oscillations are characterized by periods of enhanced and suppressed convective activity corresponding to positive and negative ocean layer temperature anomalies. The enhanced phase of the oscillations is distinguished by an abundance of organized convection. Similar to the suppressed phase of the MJO, the suppressed phase of the low frequency oscillation in the column model is characterized by

small cloud coverage and intermittent deep convection. Cloud–radiation feed-back introduces a further stochastic element to the model, as the cloud fractions are governed by a stochastic birth–death process (coupled with the large scale variables), and it thereby contributes to higher frequency variability.

In spatially extended simulations of Section 4, the radiative convective feed-back mechanism is shown to increase the mean and variability of Walker circulation. This increase in the variability is associated with low frequency convective features such as propagating synoptic scale systems originating inside of the enhanced SST area. These are similar to the cycles observed in CRM simulations (Slawinska et al., 2014b) but at a very small fraction of the computational cost. We use Fourier spectrum analysis to show that the ocean layer depth as well as the strength of the imposed SST gradient can be used to control the frequency of these low frequency convective features.

The study illustrates both the usefulness and limitation of the column model simulations for testing the mechanisms. Column simulations are essential in exploring parameter space and finding some of the catastrophic instabilities. However, the column simulations can easily exaggerate the impact of the studied mechanisms. In particular, it is clear that the intermediate and low frequency time scale features seen in the column simulations are distorted and to some extent neutralized by fast mixing by the horizontally propagating gravity waves in the spatially extended simulations.

Overall, the study develops and tests a systematic framework for incorporating parameterized radiative cloud feedback and ocean coupling. The two mechanisms are shown here to improve the representation of intraseasonal and seasonal variability. This new variability has a subtle effect on the mean state and precipitation statistics and has a great potential for interaction with other low frequency phenomena. These schemes can readily be implemented in deterministic and stochastic multcloud model GCM simulations (Khouider et al., 2011; Deng et al., 2014; Khouider et al., 2011; Ajayamohan et al., 2013, 2014), which already capture essential low frequency features, such as the MJO.

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Appendix A. Evaporative cooling at the sea surface

We recall that evaporation term E , discussed in Section 2.4 is given by

$$E = \frac{1}{\tau_e}(\theta_{eb}^*(T_s) - \theta_{eb}) \quad (\text{A.1})$$

where τ_e is evaporative time scale and θ_{eb} is the boundary layer equivalent potential temperature, and θ_{eb}^* is the saturation boundary layer equivalent potential temperature given as a function of SST. According to Claussius-Clapeyron equation, the saturation equivalent potential temperature has form

$$\theta_{eb}^* = \theta_b \exp\left(\frac{L_v q^*(T_b)}{c_p T_b}\right) \quad (\text{A.2})$$

Here T_b is the boundary layer temperature, θ_b is the boundary layer potential temperature, L_v is the latent heat of vaporization, q^* is the saturation specific humidity, and c_p is the heat capacity of dry air. Although L_v depends on temperature and c_p depends on moisture content, the variations are small and we assume that both are constants given in Table 1.

In (A.2), $q^*(T)$ is the saturation specific humidity, which for an ideal gas is given by

$$q^* = \frac{e^*}{e^* + \frac{R_v}{R_d}(p - e^*)} \quad (\text{A.3})$$

Here R_v is the gas constant for water vapor, R_d is the gas constant for dry air, p is the pressure, and e^* is the saturation water vapor pressure. I will

take R_v and R_d to have the constant values given in Table 1. I will also take
650 $p = p_s = 1000$ hPa, since the latent heat exchange takes place at the ocean
surface. If we assume that L_v is constant, then the equation (A.3) can be
integrated to give the following equation for e^* :

$$e^*(T) = e^*(T_{ref}) \exp\left(\frac{L_v}{R_v} \left[\frac{1}{T_{ref}} - \frac{1}{T}\right]\right) \quad (\text{A.4})$$

We therefore have θ_{eb}^* as a function of T by using (A.2), (A.3), and (A.4).
We approximate T_b and θ_b in these equations with the SST T_s , and we assume
655 L_v and c_p take the constant values listed in Table 1. (A.2), (A.3), and (A.4)
take the form

$$\theta_{eb}^* = \theta_b \exp\left(\frac{L_v q^*(T_s)}{c_p T_s}\right) \quad (\text{A.5})$$

$$q^*(T_s) = \frac{e^*(T_s)}{e^*(T_s) + \frac{R_v}{R_d}(p - e^*(T_s))} \quad (\text{A.6})$$

$$e^*(T_s) = e^*(T_{ref}) \exp\left(\frac{L_v}{R_v} \left[\frac{1}{T_{ref}} - \frac{1}{T_s}\right]\right) \quad (\text{A.7})$$

The equations (A.5), (A.6), and (A.7) show nonlinear relationship between θ_{eb}^*
and T_s . It appears that θ_{eb}^* rises by 5 K for 1K rise in T_s , in the range 20-35
K as seen in Figure A.13 (taking $T_{ref} = 301$ K). A linear approximation of
660 this nonlinear relationship below

$$\theta_{eb}^* \approx 5T_s \quad (\text{A.8})$$

is used in the model computations.

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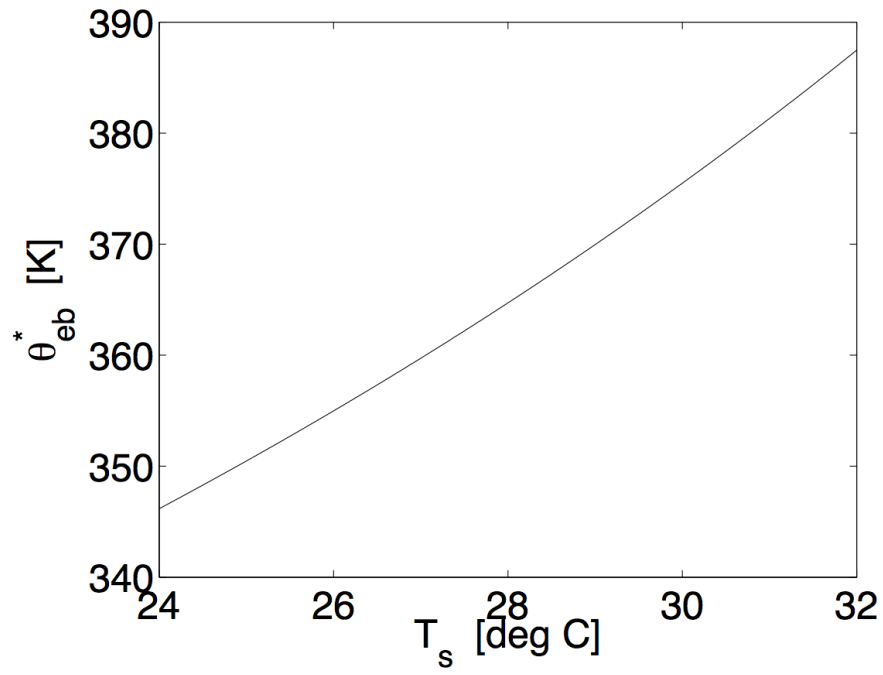


Figure A.13: The nonlinear relationship between θ_{eb}^* and T_s . It appears that θ_{eb}^* rises by 5 K for 1K rise in T_s . A linear approximation of this relationship is use in the model computations.